# Time Series Analyses in a Changing Climate Shane Elipot

The Rosenstiel School of Marine and Atmospheric Science, University of Miami

**Jonathan Lilly** 

NorthWest Research Associates, Seattle

*November 8, 2016* 

Sponsored by: the FP7 Integrated Infrastructure Initiative FixO3; the Horizon 2020 project ENVRIplus; the Arctic Marine Geology and Geophysics (AMGG) Research School

#### Foreword

This Tromsø lecture is heavily based on a longer course (The Oslo Lectures) given by Jonathan M. Lilly in Oslo at the invitation of the Norwegian Research School in Climate Dynamics (ResClim), during the week May 23–27, 2016. JML's Oslo Lectures are freely available for online viewing or download at {www.jmlilly.net/talks/oslo/index.html}

This current Tromsø lecture is available for download on my website at {selipot.github.io} under {Talks}.

The practical session this afternoon will be conducted using the {Matlab} computational software and the freely available Matlab toolbox **jLab** written by Jonathan Lilly and freely available {here}. Please download, install, and test the toolbox before the practical session.

In addition, we will be using the code and data I have made available for download in a GitHub repository {github.com/selipot/tromso}.

#### Outline

- 1. Descriptive statistics (the time domain)
- 2. Stationarity, non-stationarity, and trends
- 3. Fourier Spectral analysis
- 4. Bivariate time series
- 5. Epilogue
- 6. Practical

## 1. Descriptive Statistics (The time domain)

#### The Sample Interval

We have a sequence of N observations

$$x_n, \qquad n=0,1,2,\ldots N-1$$

which coincide with times

$$t_n, \qquad n=0,1,2,\ldots N-1.$$

The sequence  $x_n$  is called a discrete time series.

It is assumed that the *sample interval*,  $\Delta_t$ , is constant

$$t_n = n \Delta_t$$

with the time at n=0 defined to be 0. The duration is  $T=N\Delta_t$  .

If the sample interval in your data is not uniform, the first processing step is to interpolate it to be so.

## The Underlying Process

A critical assumption is that there exists some "process" x(t) that our data sequence  $x_n$  is a sample of:

$$x_n=x(n\Delta_t), \qquad n=0,1,2,\ldots N-1.$$

Unlike  $x_n$ , x(t) is believed to exist for *all times*.

- (i) The process x(t) exists in continuous time, while  $x_n$  only exists at discrete times.
- (ii) The process x(t) exists for all past and future times, while  $x_n$  is only available over a certain time interval.

It is the properties of x(t) that we are trying to estimate, based on the available sample  $x_n$ .

#### Measurement Noise

In reality, the measurement device and/or data processing probably introduces some artifical variability, termed *noise*.

It is more realistic to consider that the observations  $x_n$  contain samples of the process of interest, y(t), plus some noise  $\epsilon_n$ :

$$x_n = y(n\Delta_t) + \epsilon_n, \qquad n = 0, 1, 2, \ldots N-1.$$

This is an example of the *unobserved components model*. This means that we *believe* that the data is composed of *different components*, but we cannot observe these components individually.

The process y(t) is potentially obscured or degraded by the limitations of data collection in three ways: (i) finite sample interval, (ii) finite duration, (iii) noise.

Because of this, the time series is an *imperfect* representation of the real-world processes we are trying to study.

#### Time versus Frequency

There are two complementary (opposite) points of view regarding the time series  $x_n$ .

The first regards  $x_n$  as being built up as a sequence of discrete values  $x_0, x_2, \ldots x_{N-1}$ .

This is the domain of *statistics*: the mean, variance, histogram, etc.

When we look at data statistics, generally, the order in which the values are observed *doesn't matter*.

The second point of view regards  $x_n$  as being built up of sinusoids: purely periodic functions spanning the whole duration of the data.

This is the domain of Fourier spectral analysis.

In between these two extremes is wavelet analysis which is not covered here, see the Oslo lectures.



#### Time-Domain Statistics

A good place to start is with the very simplest tools.

The sample mean describes a "typical" value:

$$\hat{\mu}_x \equiv rac{1}{N} \sum_{n=0}^{N-1} x_n$$

The *sample variance* gives the spread about the mean:

$$\hat{\sigma}_x^2 \equiv rac{1}{N} \sum_{n=0}^{N-1} (x_n - \hat{\mu}_x)^2$$

"Sample" here means that it is computed from the observed data, as opposed to being a property of the assumed underlying process x(t). That's why we use  $\hat{\cdot}$ .

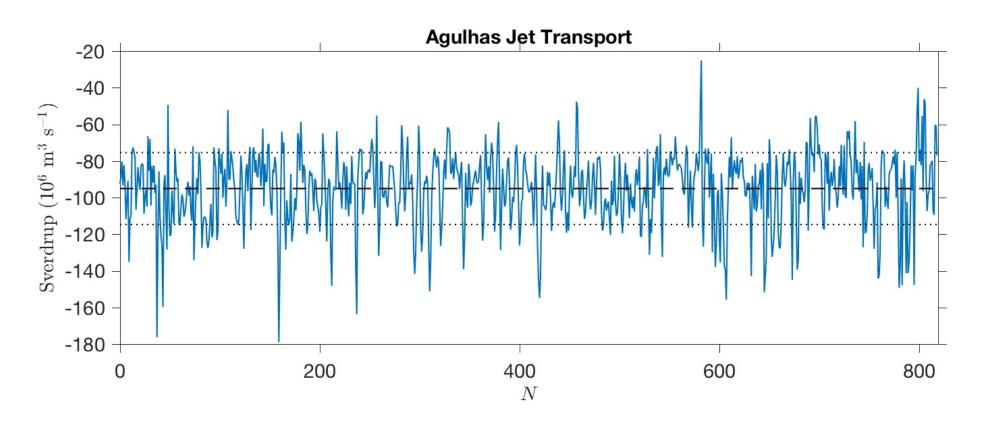
The mean and variance are called the first two *moments* of the distribution of values associated with the process.

#### Example

22-year time series of Agulhas current transport:

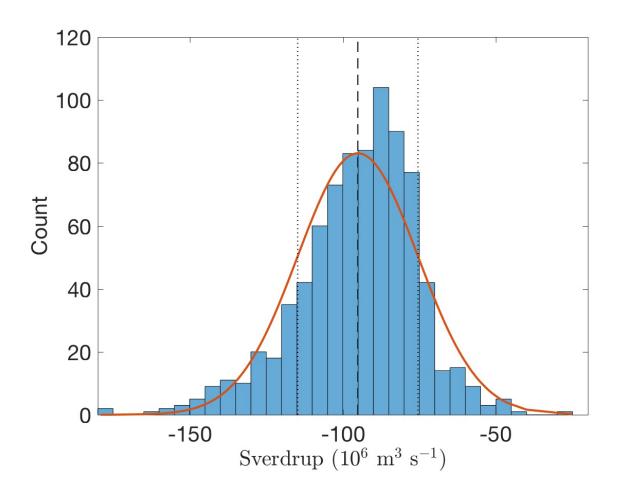
Sample mean  $\hat{\mu}_x$  is -95 Sverdrup (1 Sv =  $10^6$  m<sup>3</sup> s<sup>-1</sup>)

Sample standard deviation  $\sqrt{\hat{\sigma}_x^2}$  is 20 Sv



## Histogram

The mean, variance, skewness, and kurtosis (seen next) describe aspects of the *histogram*: the observed distribution of data values.

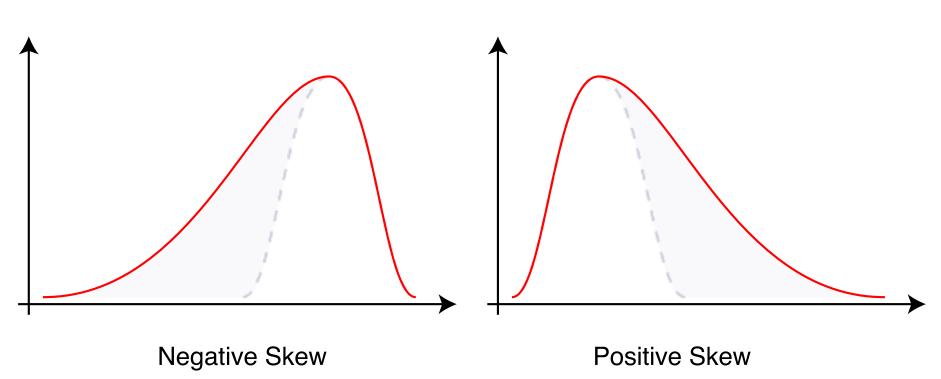


Here is the histogram of *all* Agulhas Jet transport (blue), versus Gaussian noise having the same mean and variance (orange).

#### Skewness

The *skewness* describes the tendency for an *asymmetry* between positive excursions and negative excursions:

$$\hat{\gamma}_x \equiv rac{1}{\hat{\sigma}_x^3} rac{1}{N} \sum_{n=0}^{N-1} (x_n - \hat{\mu}_x)^3$$



#### Kurtosis

The *kurtosis* is said to either measure *peakedness* (concentration near  $\hat{\mu}_x$ ), or a tendency for *long tails* (concentration far from  $\hat{\mu}_x$ ):

$$\hat{\kappa}_x \equiv rac{1}{\hat{\sigma}_x^4} rac{1}{N} \sum_{n=0}^{N-1} (x_n - \hat{\mu}_x)^4$$

Actually, it measures both. Kurtosis is a measure of the spread of  $x_n$  about the *two points*  $\mu_x \pm \sigma_x$ . This can happen *either* for peakness or for long tails! See Moors (1986), "The Meaning of Kurtosis: Darlington Reexamined".

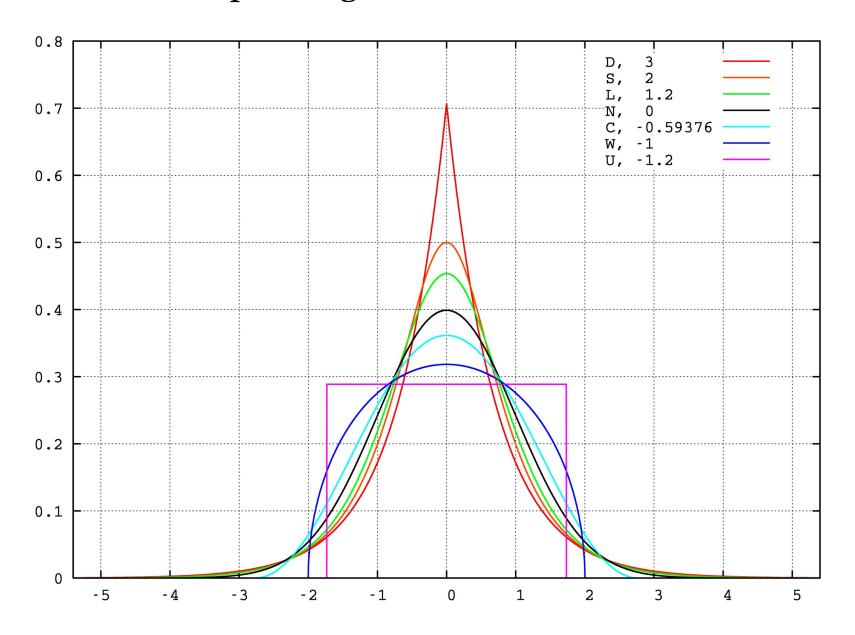
Because the value of kurtosis for a Gaussian process can be shown to be equal to 3, one sometimes encounters the *excess kurtosis* 

$$ilde{\kappa}_x \equiv \kappa_x - 3.$$

Values of  $\tilde{\kappa}_x > 0$  mean the data is more kurtotic—peaked or long-tailed—than a Gaussian, while  $\tilde{\kappa}_x < 0$  means it is less so.

#### Illustration of Kurtosis

Distributions corresponding to different values of excess kurtosis.

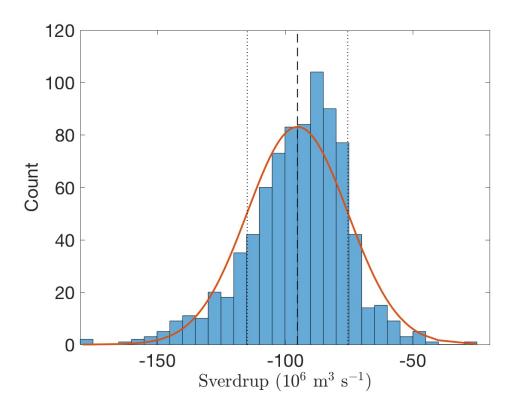


Positive excess kurtosis corresponds to long tails and peakedness.



## Example

$$\hat{\gamma}_x-=0.6,\quad \hat{\kappa}_x-3=1.14\,.$$



Here is the histogram of *all* Agulhas Jet transport (blue), versus Gaussian noise having the same mean and variance (orange).

#### Process stats vs Estimates

How can we be sure that the *sample* statistics for  $x_n$  are the statistics of the process x(t)? We can't, but we can estimate an uncertainty for our estimates, also called an *error*.

As an example recall the sample mean is  $\hat{\mu}_x \equiv \frac{1}{N} \sum_{n=0}^{N-1} x_n$ .

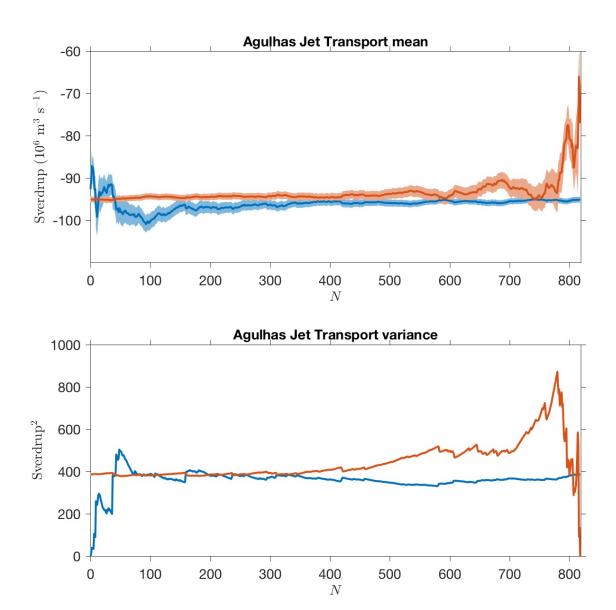
It can be shown that the theoretical variance of the sample mean is

$$\sigma_{\hat{\mu}_x}^2 \equiv rac{\sigma_x^2}{N}.$$

so that **estimates** of the *standard error of the mean* can be calculated as  $\frac{\hat{\sigma}_x}{\sqrt{N}}$ . This is a general principle: abstract concepts or mathematical objects are estimated by *estimators*, which have errors (bias and variance).

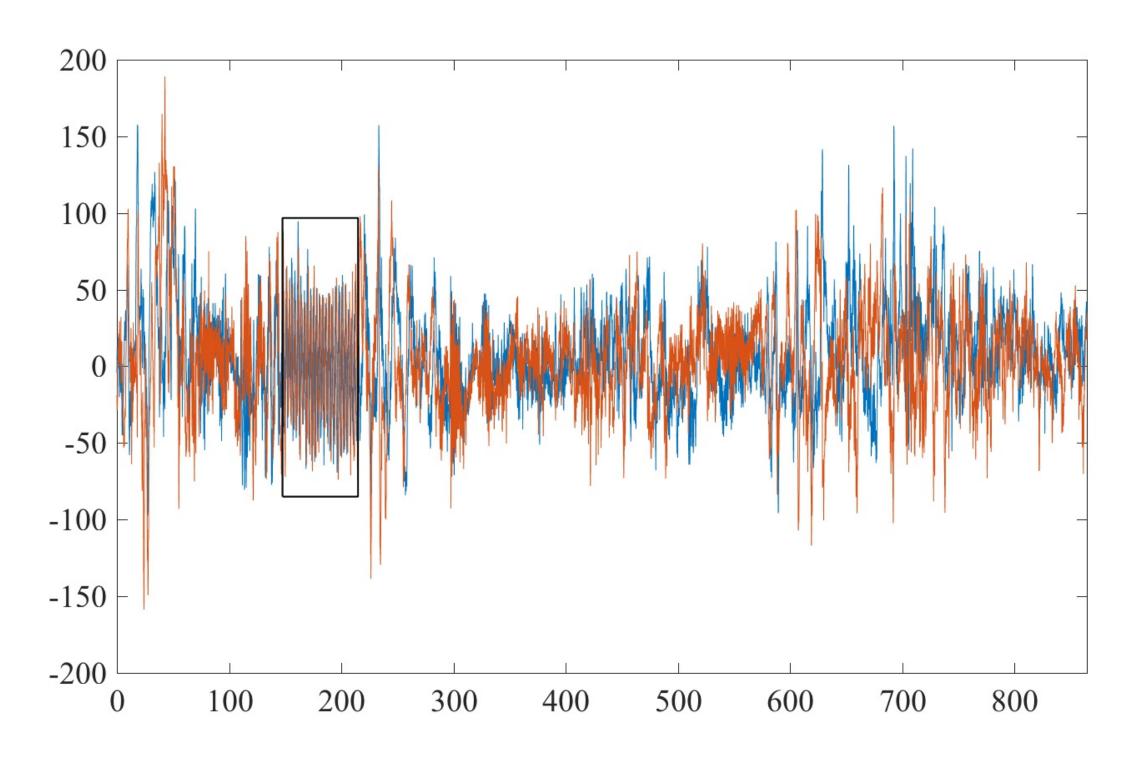
#### Example

Plots of  $\hat{\mu}_x \pm \hat{\sigma}_{\mu_x}$  and  $\hat{\sigma}_x^2$  as a function of increasing N in chronological order (blue) or reverse chronological order (orange).



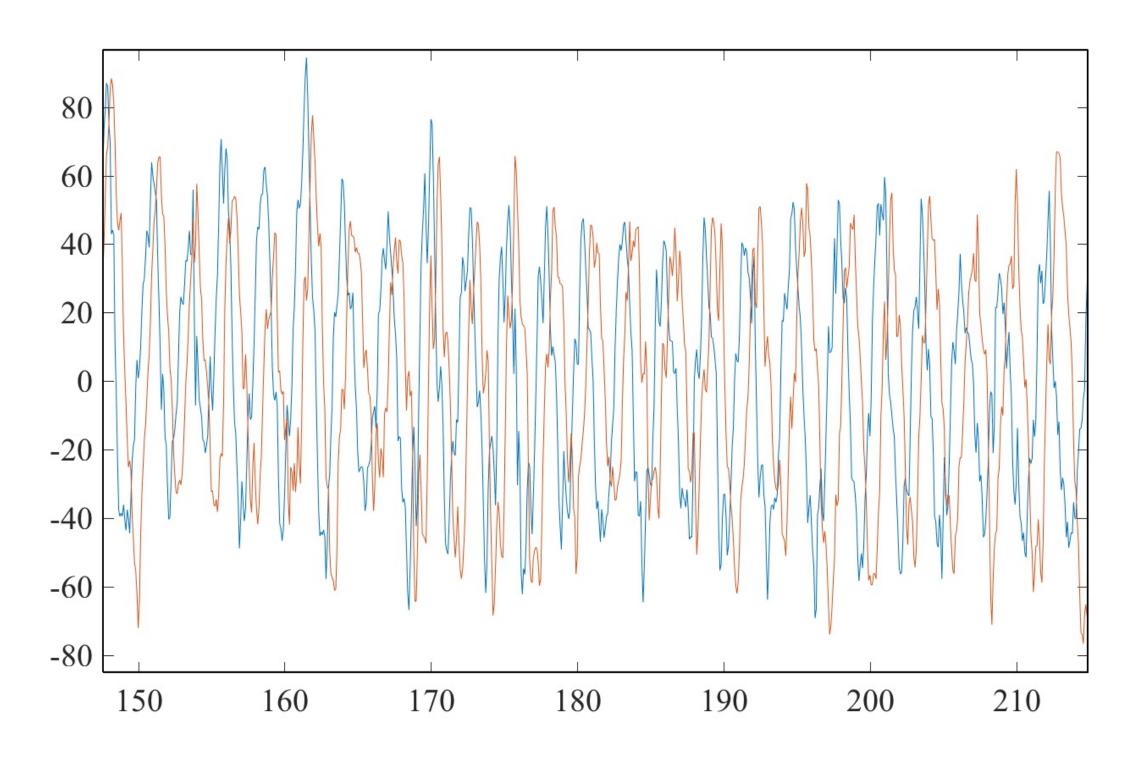
## 2. Stationarity vs nonstationarity, trends

## First Example



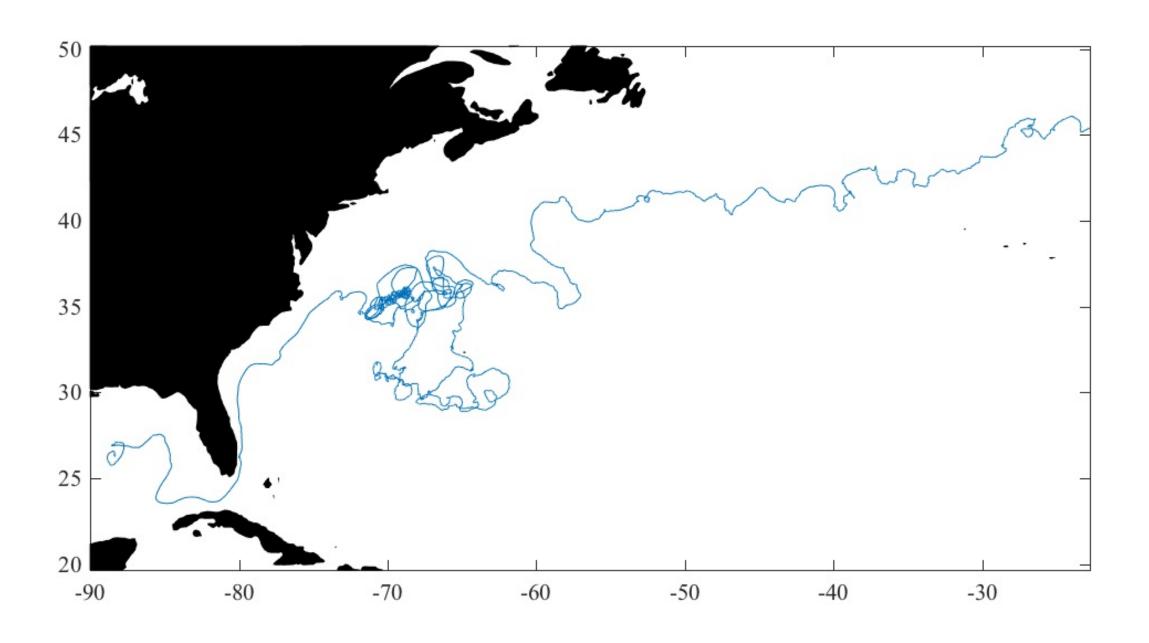


## First Example





## First Example





#### Observable Features

- 1. The data consists of two time series that are similar in character.
- 2. Both time series present a superposition of scales and a high degree of roughness.
- 3. The data seems to consist of different time periods with distinct statistical characteristics—the data is *nonstationary*.
- 4. Zooming in to one particular period show regular oscillations of roughly uniform amplitude and frequency.
- 5. The phasing of these show a circular polarization orbited in a counterclockwise direction.
- 6. The zoomed-in plot shows a fair amount of what appears to be measurement noise superimposed on the oscillatory signal.

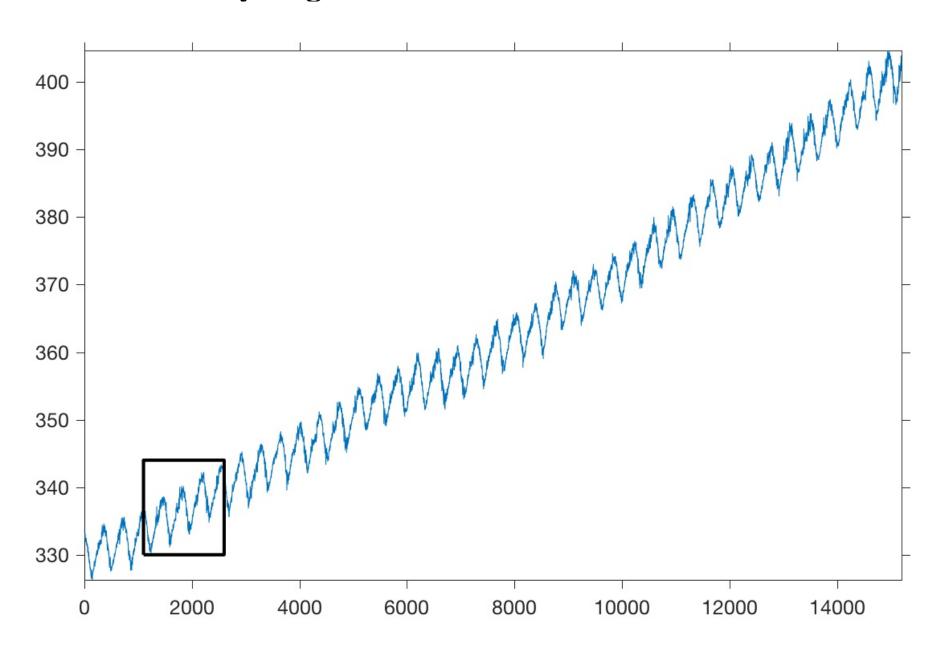
#### Observable Features

- 1. The data consists of two time series that are similar in character.
- 2. Both time series present a superposition of scales and a high degree of roughness.
- 3. The data seems to consist of different time periods with distinct statistical characteristics—the data is *nonstationary*.
- 4. Zooming in to one particular period show regular oscillations of roughly uniform amplitude and frequency.
- 5. The phasing of these show a circular polarization orbited in a counterclockwise direction.
- 6. The zoomed-in plot shows a fair amount of what appears to be measurement noise superimposed on the oscillatory signal.

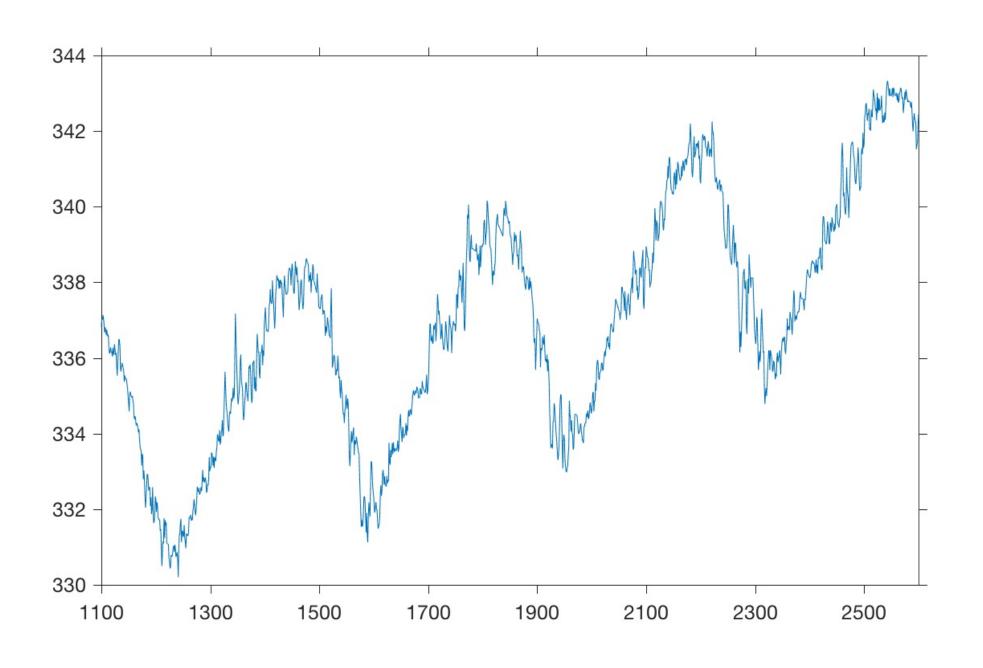
This is a record of velocities of a single surface drifter at 6-hour intervals. All Surface drifter data are freely available from the Data Assembly Center of the Global Drifter Program {www.aoml.noaa.gov/phod/dac/}.

## Second Example

Can you guess what this time series is?



## Second Example





#### Observable Features

- 1. The data exhibit a very strong positive trend, roughly linear with time. Thus, this time series does not present a **mean statistics** that represents a "typical" value.
- 2. On top of the trend there seems to be a sinusoid-like oscillation that does not appear to change with time.
- 3. The zoomed-in plot shows noise superimposed on the sinusoidal and trend processes.

#### Observable Features

- 1. The data exhibit a very strong positive trend, roughly linear with time. Thus, this time series does not present a **mean statistics** that represents a "typical" value.
- 2. On top of the trend there seems to be a sinusoid-like oscillation that does not appear to change with time.
- 3. The zoomed-in plot shows noise superimposed on the sinusoidal and trend processes.

This is a record of daily atmospheric CO<sub>2</sub> measured at Mauna Loa in Hawaii at an altitude of 3400 m. Data from Dr. Pieter Tans, NOAA/ESRL ({www.esrl.noaa.gov/gmd/ccgg/trends/}) and Dr. Ralph Keeling, Scripps Institution of Oceanography ({scrippsco2.ucsd.edu}).

We will investigate further this time series during the practical session this afternoon.

#### Non-stationarity

The sample statistics may be changing with time because the underlying process (that is its statistics) is changing with time. The process is said to be "non-stationary".

Sometimes we need to re-think our model for the underlying process. Let us hypothesized that the process x(t) is the sum of an unknown process y(t), plus a linear trend a, plus noise:

$$x(t) = y(t) + at + \epsilon(t),$$

#### Non-stationarity

The sample statistics may be changing with time because the underlying process (that is its statistics) is changing with time. The process is said to be "non-stationary".

Sometimes we need to re-think our model for the underlying process. Let us hypothesized that the process x(t) is the sum of an unknown process y(t), plus a linear trend a, plus noise:

$$x(t) = y(t) + at + \epsilon(t),$$

or maybe the trend is better described as being quadratic with time because of an acceleration:

$$x(t) = y(t) + bt^2 + at + \epsilon(t).$$

#### Non-stationarity

The sample statistics may be changing with time because the underlying process (that is its statistics) is changing with time. The process is said to be "non-stationary".

Sometimes we need to re-think our model for the underlying process. Let us hypothesized that the process x(t) is the sum of an unknown process y(t), plus a linear trend a, plus noise:

$$x(t) = y(t) + at + \epsilon(t),$$

or maybe the trend is better described as being quadratic with time because of an acceleration:

$$x(t) = y(t) + bt^2 + at + \epsilon(t).$$

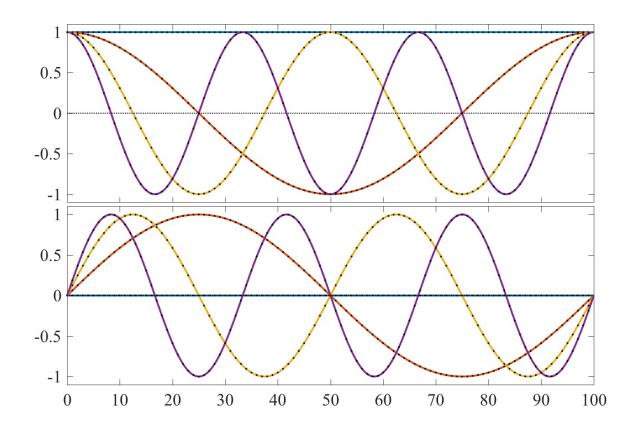
The goal is then to estimate the unknowns a, b, which consists of methods of analyses generally called "parametric". It is a bit like analyzing the data in terms of its statistics (with no prior expectations) or assuming a form for the data.

## 3. Fourier Spectral Analysis

#### Complex Fourier Series

It is possible to represent a discrete time series  $x_n$  as a sum of complex exponentials, a complex Fourier series:

$$x_n = rac{1}{N\Delta_t} \sum_{m=0}^{N-1} X_m e^{i2\pi m n/N}, \qquad n = 0, 1, \ldots N-1$$



We leave out for now how to obtain the complex coefficients  $X_m$  ...



## **About Frequency**

You will typically find two frequency notations:

$$\cos(2\pi f t)$$
 or  $\cos(\omega t)$ 

f is called the cyclic frequency. Its units are cycles/time. Example: Hz = cycles/sec.

 $\omega=2\pi f$  is called the *radian* or *angular* frequency. Its units are rad/time. The associated *period* of oscillation is  $P=1/f=2\pi/\omega$ .

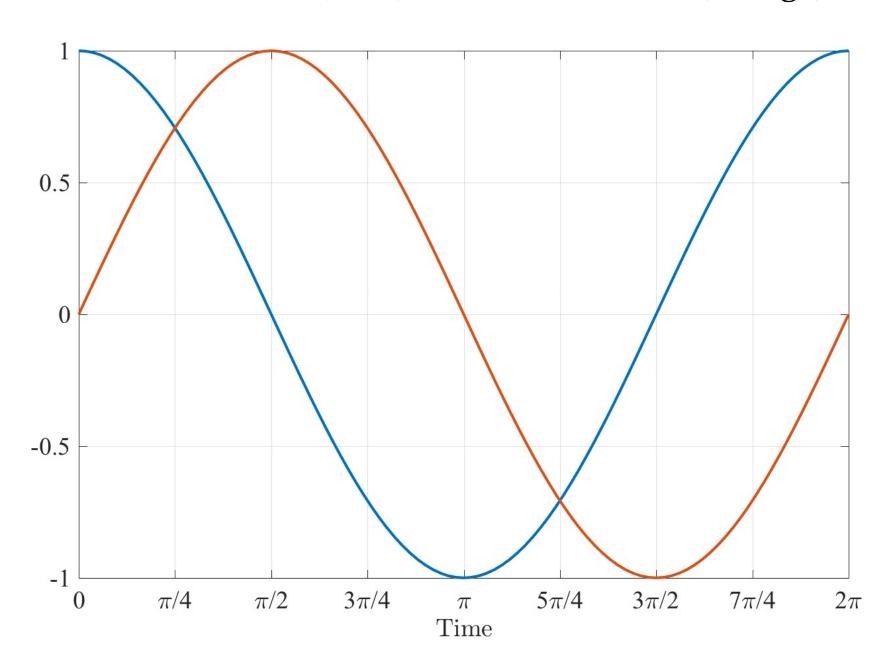
As t goes from 0 to  $1/f=2\pi/\omega=P$ ,  $2\pi ft$  goes from 0 to  $2\pi$  and  $\omega t$  goes from 0 to  $2\pi$ .

A very common error in Fourier analysis is mixing up cyclic and radian frequencies!

Note: neither "cycles" nor "radians" actually have any units, thus both f and  $\omega$  have units of 1/time. However, specifying for example 'cycles per day' or 'radians per day' helps to avoid confusion.

#### Review: Sinusoids

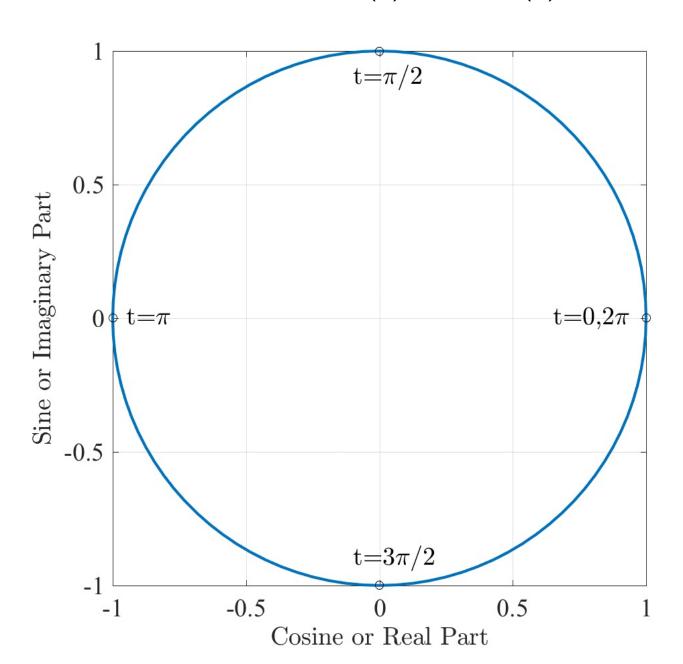
Cosine function (blue) and sine function (orange)



## Complex Exponentials, 2D

Now consider a plot cos(t) vs. sin(t).

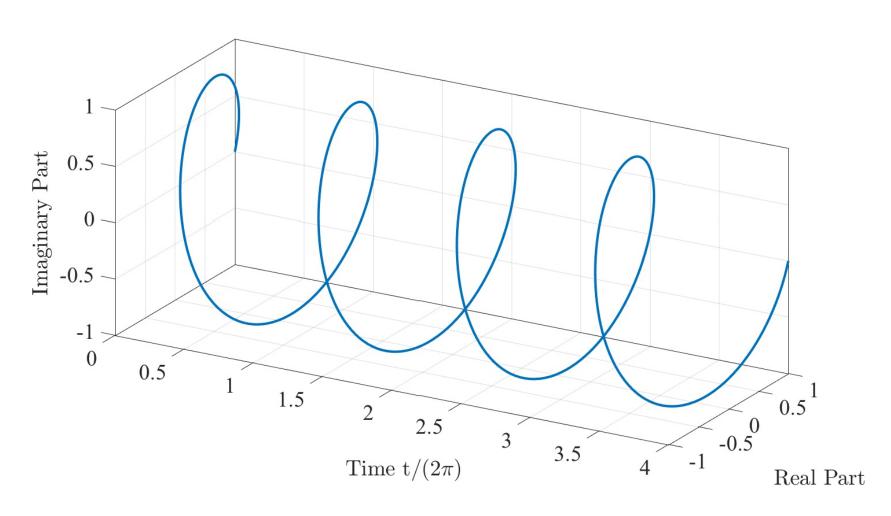
That's the same as  $\cos(t) + i\sin(t) = e^{it}$  .



## Complex Exponentials, 3D

This is better seen in 3D as a spiral as time increases.

$$\cos(t) + i\sin(t) = e^{it}$$



#### The complex Fourier series

The discrete time series  $x_n$  can written as a sum of complex exponentials:

$$x_n = rac{1}{N\Delta_t} \sum_{m=0}^{N-1} X_m e^{i2\pi m n/N} = rac{1}{N\Delta_t} \sum_{m=0}^{N-1} X_m e^{i2\pi n \Delta_t \cdot (m/N\Delta_t)}, \quad n = 0, 1, \dots$$



#### The complex Fourier series

The discrete time series  $x_n$  can written as a sum of complex exponentials:

$$x_n = rac{1}{N\Delta_t} \sum_{m=0}^{N-1} X_m e^{i2\pi m n/N} = rac{1}{N\Delta_t} \sum_{m=0}^{N-1} X_m e^{i2\pi n \Delta_t \cdot (m/N\Delta_t)}, \quad n = 0, 1, \dots$$

The mth term behaves as  $e^{i2\pi f_m n\Delta_t}=\cos(2\pi f_m n\Delta_t)+i\sin(2\pi f_m n\Delta_t)$  , where  $f_m\equiv m/N\Delta_t$  .

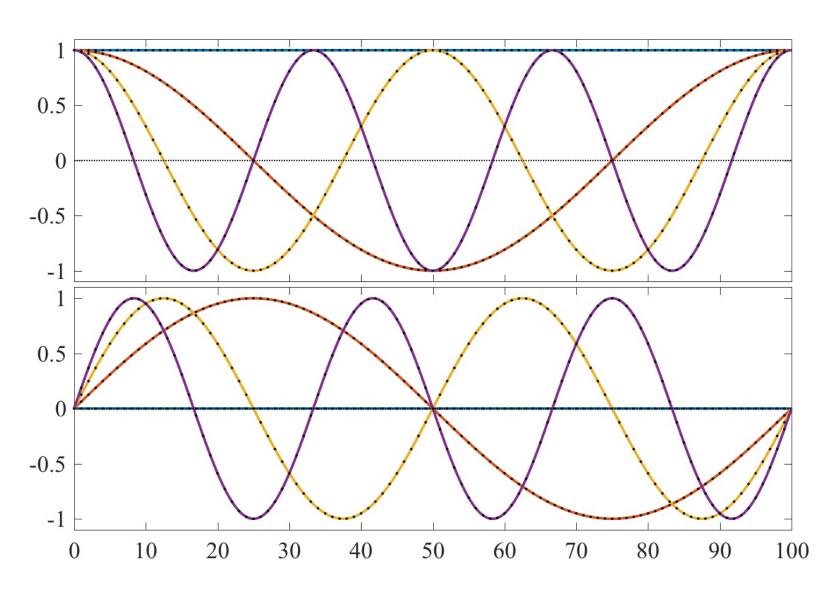
Note that in the literature,  $\Delta_t$  is often set to one, and thus omitted, leading to a lot of confusion (including for me!).

The quantity  $f_m \equiv m/N\Delta_t$  is called the mth Fourier frequency. The period associated with  $f_m$  is  $1/f_m = \Delta_t N/m$ . Thus m tells us the number of oscillations contained in the length  $N\Delta_t$  time series.

#### Continuous Time

 $\cos(2\pi f_m t)$  and  $\sin(2\pi f_m t)$ 

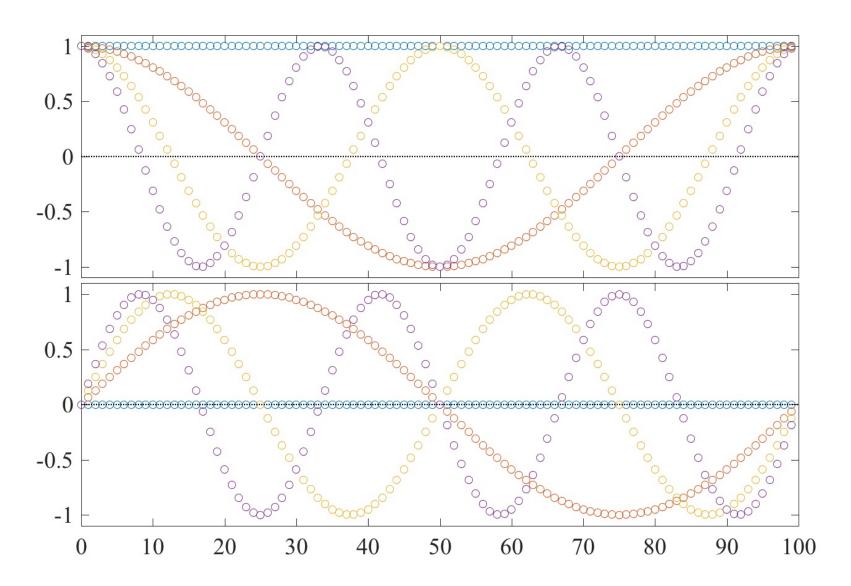
$$f_m=0,\,1/100,\,2/100,\,3/100 \qquad t=[0\dots 100]$$



#### Discrete Time

 $\cos(2\pi f_m n \Delta_t)$  and  $\sin(2\pi f_m n \Delta_t)$ 

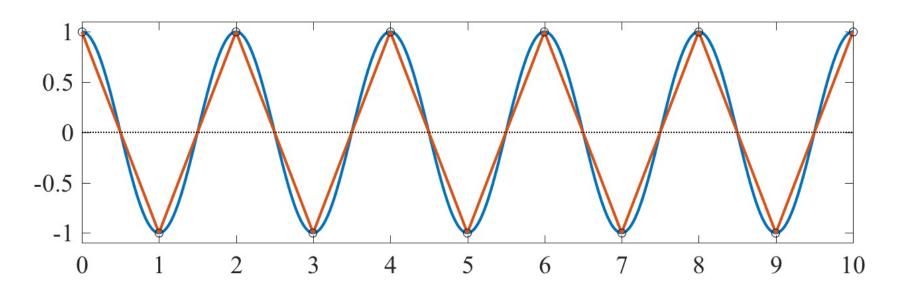
$$f_m=0,\,1/100,\,2/100,\,3/100$$
  $n=0,1,2,\dots 99$   $\Delta_t=1$ 



## The Nyquist Frequency

The single most important frequency is the *highest resolvable* frequency, the *Nyquist frequency*.

$$f^{\mathcal{N}} \equiv rac{1}{2\Delta_t} = rac{1}{2} \cdot rac{1}{\Delta_t} \qquad \omega^{\mathcal{N}} \equiv rac{1}{2} \cdot rac{2\pi}{\Delta_t} = rac{\pi}{\Delta_t}$$



The highest resolvable frequency is *half* the *sampling rate* or *one cycle per two sampling intervals*.

$$e^{i2\pi f^{\mathcal{N}} n\Delta_t} = e^{i2\pi\cdot 1/(2\Delta_t)\cdot n\Delta_t} = e^{i\pi n} = (-1)^n = 1, -1, 1, -1, \ldots$$

Note that there is no "sine" component at Nyquist in the Fourier series!

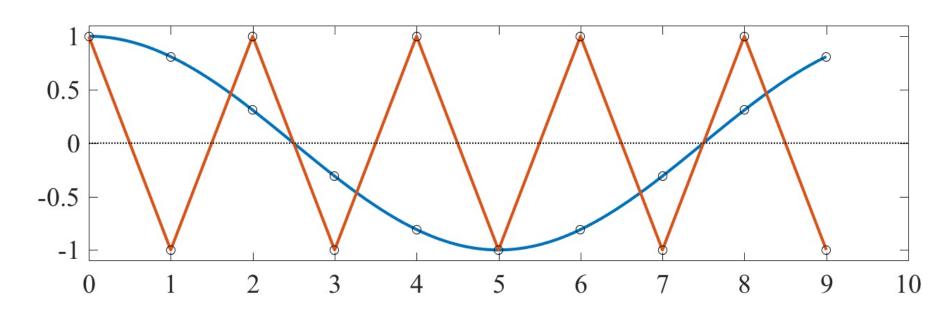
36 / 102



## The Rayleigh Frequency

The second most important frequency is the *lowest resolvable* frequency, the *Rayleigh frequency*.

$$f^{\mathcal{R}} \equiv rac{1}{N\Delta_t} \qquad \omega^{\mathcal{R}} \equiv rac{2\pi}{N\Delta_t}$$



The lowest resolvable frequency is one cycle over the *entire record*. Here the sample interval  $\Delta_t=1$  and the number of points is N=10.

#### Importance of Rayleigh

The Rayleigh frequency  $f^{\mathcal{R}}$  is important because it gives the *spacing between* the Fourier frequencies:

$$f_0=0$$
 ,  $f_1=rac{1}{N\Delta_t}$  ,  $f_2=rac{2}{N\Delta_t},\ldots$   $f_n=n\,f^{\mathcal{R}},$   $f^{\mathcal{R}}=rac{1}{N\Delta_t}$ 

Thus, it controls the *frequency-domain resolution*. If you want to distiguish between two closely spaced peaks, you need the dataset *duration* to be sufficiently *large* so that the Rayleigh frequency is sufficiently *small*.

#### Importance of Rayleigh

The Rayleigh frequency  $f^{\mathcal{R}}$  is important because it gives the *spacing between* the Fourier frequencies:

$$f_0=0$$
 ,  $f_1=rac{1}{N\Delta_t}$  ,  $f_2=rac{2}{N\Delta_t},\ldots$   $f_n=n\,f^{\mathcal{R}},$   $f^{\mathcal{R}}=rac{1}{N\Delta_t}$ 

Thus, it controls the *frequency-domain resolution*. If you want to distiguish between two closely spaced peaks, you need the dataset *duration* to be sufficiently *large* so that the Rayleigh frequency is sufficiently *small*.

As an example, the two principal semi-diurnal tidal "species" have period of 12 h  $(M_2)$  and 12.4206012 h  $(S_2)$ . The minimum record length to distinguish the two frequencies is thus

$$N\Delta_t = rac{1}{f^{\mathcal{R}}} = rac{1}{f^{S_2} - f^{M_2}} = rac{1}{rac{1}{12} - rac{1}{12.4206012}} = 354.36 ext{ hours.}$$

# Rayleigh and Nyquist frequencies

The ratio of the Rayleigh to Nyquist frequencies tells you how many different frequencies you can resolve.

$$rac{f^{\mathcal{N}}}{f^{\mathcal{R}}} = rac{N\Delta_t}{2\Delta_t} = rac{N}{2}$$

So why do we have N frequencies in the sum for the complex Fourier series?

$$x_n = rac{1}{N\Delta_t}\sum_{m=0}^{N-1} X_m e^{i2\pi mn/N}$$

#### The Fourier Frequencies

The first few Fourier frequencies  $f_m = m/(N\Delta_t)$  are:

$$f_0=rac{0}{N\Delta_t},\quad f_1=rac{1}{N\Delta_t},\quad f_2=rac{2}{N\Delta_t},\ldots$$

while the last few are

$$\dots, f_{N-2} = rac{N-2}{N\Delta_t} = rac{1}{\Delta_t} - rac{2}{N\Delta_t}, \quad f_{N-1} = rac{N-1}{N\Delta_t} = rac{1}{\Delta_t} - rac{1}{N\Delta_t}.$$

But notice that the last Fourier exponential term is

$$e^{i2\pi f_{N-1}n\Delta_t}=e^{i2\pi(N-1)n/N}=e^{i2\pi n}e^{-i2\pi n/N}=e^{-i2\pi n/N}=e^{-i2\pi f_1n\Delta_t}$$

because  $e^{i2\pi n}=1$  for all integers n! Frequencies higher than the Nyquist cannot appear due to our sample rate. Therefore, these terms instead specify terms that have a frequency less than the Nyquist but that rotate in the negative direction.

## The Fourier Frequencies

In the vicinity of m=N/2, for even N, we have

$$f_{N/2-1} = rac{1}{2\Delta_t} - rac{1}{N\Delta_t}, \quad f_{N/2} = rac{1}{2\Delta_t}, \quad f_{N/2+1} = rac{1}{2\Delta_t} + rac{1}{N\Delta_t}, \ldots$$

but actually the first frequency higher than the Nyquist is the *highest negative frequency*:

$$f_{N/2-1}=rac{1}{2\Delta_t}-rac{1}{N\Delta_t},\quad f_{N/2}=rac{1}{2\Delta_t},\ldots$$

$$f_{N/2+1}=-f_{N/2-1}=-\left(rac{1}{2\Delta_t}-rac{1}{N\Delta_t}
ight),\ldots.$$

Thus the positive frequencies and negative frequencies are both increasing *toward the middle* of the Fourier transform array.

For this reason Matlab provides fftshift, to shifts the zero frequency, *not* the Nyquist, to be in the middle of the array.



#### One-Sided vs. Two-Sided

There exists two strictly equivalent representations, *two-sided* and *one-sided*, of the discrete Fourier transform:

$$x_n=rac{1}{N\Delta_t}\sum_{m=0}^{N-1}X_me^{i2\pi mn/N},$$

$$x_n = rac{1}{N\Delta_t} X_0 + rac{2}{N\Delta_t} \sum_{m=1}^{N/2-1} A_m \cos(2\pi m n/N + \Phi_m) + X_{N/2} (-1)^n,$$

where  $A_m$  and  $\Phi_m$  are an *amplitude* and *phase*, with  $X_m = A_m e^{i\Phi_m}$  .

The two-sided representation is more compact mathematically.

For real-valued  $x_n$ , the one-sided representation is more intuitive as it expresses  $x_n$  as a sum of *phase-shifted* cosinusoids.

#### One-Sided vs. Two-Sided

There exists two strictly equivalent representations, *two-sided* and *one-sided*, of the discrete Fourier transform:

$$x_n=rac{1}{N\Delta_t}\sum_{m=0}^{N-1}X_me^{i2\pi mn/N},$$

$$x_n = rac{1}{N\Delta_t} X_0 + rac{2}{N\Delta_t} \sum_{m=1}^{N/2-1} A_m \cos(2\pi m n/N + \Phi_m) + X_{N/2} (-1)^n,$$

where  $A_m$  and  $\Phi_m$  are an *amplitude* and *phase*, with  $X_m = A_m e^{i\Phi_m}$  .

The two-sided representation is more compact mathematically.

For real-valued  $x_n$ , the one-sided representation is more intuitive as it expresses  $x_n$  as a sum of *phase-shifted* cosinusoids. A price of the one-sided form is that even and odd N are somewhat different! The expression above is for even-valued N.



#### The Forward DFT

So? How do we know the values of the Fourier coefficients  $X_m$ ? It can be shown that:

$$X_m = \Delta_t \sum_{n=0}^{N-1} x_n e^{-i2\pi m n/N}$$

This is called the discrete Fourier transform of  $x_n$ .

The DFT transforms  $x_n$  from the time domain to the frequency domain. The DFT defines a sequence of N complex-valued numbers,  $X_m$ , for  $m=0,1,2,\ldots N-1$ , which are termed the Fourier coefficients.

In Matlab, the discrete Fourier transform defined above is computed by  $fft(x) \times \Delta_t$ .

#### The Inverse DFT

In fact,

$$x_n \equiv rac{1}{N\Delta_t} \sum_{m=0}^{N-1} X_m e^{i2\pi m n/N}$$

is called the *inverse discrete Fourier transform*. It expresses how  $x_n$  may be *constructed* using the Fourier coefficients multiplying complex exponentials—or, as we saw earlier, phase-shifted sinusoids.

#### The Spectrum

One of several definitions of the spectrum, or spectral density function, at frequency  $f_m$ , is:

$$S(f_m) \equiv \lim_{N o \infty} E \left\{ rac{|X_m|^2}{N} 
ight\}.$$

 $E\{\cdot\}$  is called the *expectation*, it is a conceptual "average" over a statistical ensemble, and it cannot obtained in practice.

#### The Spectrum

One of several definitions of the spectrum, or spectral density function, at frequency  $f_m$ , is:

$$S(f_m) \equiv \lim_{N o \infty} E \left\{ rac{|X_m|^2}{N} 
ight\}.$$

 $E\{\cdot\}$  is called the *expectation*, it is a conceptual "average" over a statistical ensemble, and it cannot obtained in practice.

Formally, the function S is defined for all frequencies f, not only  $f_m$ , but as  $N \to \infty$ , the Rayleigh frequency  $1/N\Delta_t$  becomes infinitesimally small, and all frequencies are obtained.

#### The Spectrum

One of several definitions of the spectrum, or spectral density function, at frequency  $f_m$ , is:

$$S(f_m) \equiv \lim_{N o \infty} E \left\{ rac{|X_m|^2}{N} 
ight\}.$$

 $E\{\cdot\}$  is called the *expectation*, it is a conceptual "average" over a statistical ensemble, and it cannot obtained in practice.

Formally, the function S is defined for all frequencies f, not only  $f_m$ , but as  $N \to \infty$ , the Rayleigh frequency  $1/N\Delta_t$  becomes infinitesimally small, and all frequencies are obtained.

However,  $N \to \infty$  is not achievable ... Therefore, one aspect of spectral analysis is to find an acceptable **estimate** of the true, unknown, spectrum S(f) of the process x(t).

#### The Parseval Theorem

A very important theorem is *Parseval's theorem* which takes the following form for the discrete case:

$$|\Delta_t \sum_{n=0}^{N-1} |x_n|^2 = rac{1}{N\Delta_t} \sum_{m=0}^{N-1} |X_m|^2.$$

When  $\hat{\mu}_x = 0$ , this theorem shows that the total **variance** of  $x_n$  is recoverable from the sum of absolute Fourier coefficients squared.

Which can be interpreted as saying that the spectrum gives you the distribution of the variance as a function of frequency.



#### Spectral Estimates

The simplest way to estimate the *spectrum* S(f) function of frequency f is to simply take the modulus squared of the Fourier transform,

$$\widehat{S}(f_m) = \widehat{S}_m \equiv rac{1}{N} |X_m|^2, \qquad m = 0, 1, 2, \ldots, (N-1).$$

This quantity is known as the *periodogram*.

Note that the Matlab fft(x) command assumes  $\Delta_t = 1$  so you need to plot abs(fft(x))<sup>2</sup> ×  $\Delta_t/N$ .

As we shall see, the periodogram is *not* the spectrum! It is an *estimate* of the spectrum—and generally speaking, a very poor one.

It is also said to be the *naive* spectral estimate, meaning it is the spectral estimate that you get if you don't know that there is something better. Please do not use the periodogram in your publications.

## The Multitaper Method

An alternate spectral estimate called the *multitaper* method. Here is a quick sketch of this method.

We form a set of K different sequences the same length as the data, that is, having N points in time. These sequences are chosen from a special family of functions that is closely related to familiar orthogonal functions, e.g. the Hermite functions.

These K different sequences are denoted as  $\psi_n^{\{k\}}$  for  $k=1,2,\ldots K$  . For each of these sequence, we form a spectral estimate as

$$\widehat{S}_m^{\{k\}} \equiv \left| \Delta_t \sum_{n=0}^{N-1} \psi_n x_n \, e^{-i2\pi m n/N} 
ight|^2, \qquad n = 0, 1, 2, \ldots, (N-1).$$

which involves *multiplying* the data by the sequence  $\psi_n^{\{k\}}$  before taking the Fourier transform.

#### The Multitaper Method

The action of multiplying the data by some sequence before Fourier transforming it, as in

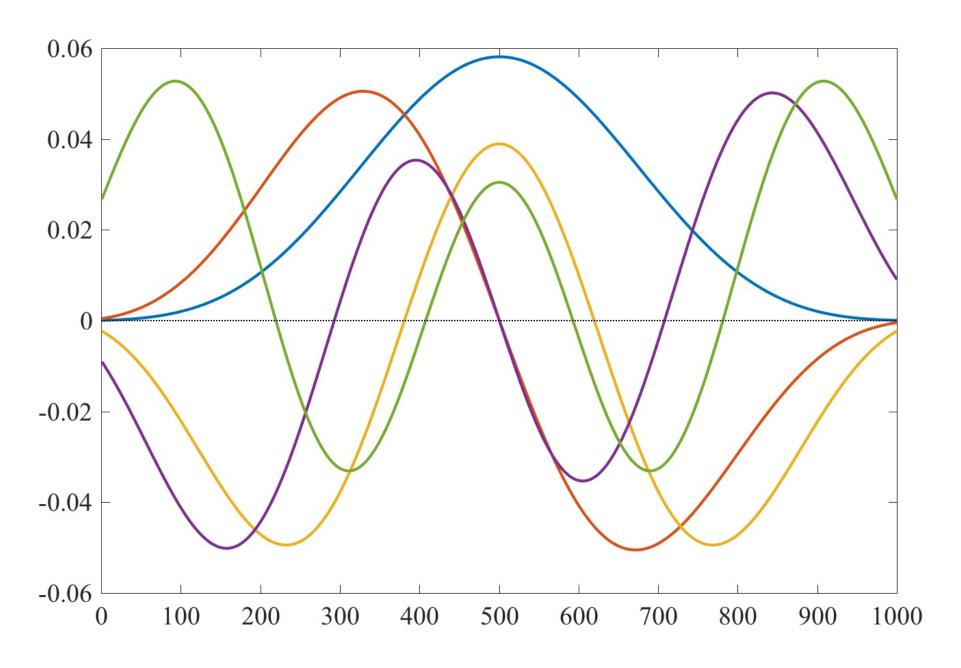
$$\widehat{S}_m^{\{k\}} \equiv \left| \Delta_t \sum_{n=0}^{N-1} \psi_n x_n \ e^{-i2\pi m n/N} 
ight|^2, \qquad n = 0, 1, 2, \ldots, (N-1)$$

is called *tapering*. The goal is to reduce the **bias** (systematic error) of the spectral estimate. These K different individual estimates (aka *eigenspectra*), are combined into one *average* spectral estimate, in order to reduce the **variance** (random error) of the estimate

$$\widehat{S}_m^\psi \equiv rac{1}{K} \sum_{k=1}^K \widehat{S}_m^{\{k\}}.$$

The multitaper method therefore involves two steps: (i) *tapering* the data, and (ii) *averaging* over multiple individual spectral estimates.

#### The Taper Functions



Here K=5 Slepian tapers are shown. These are orthogonal functions that become more oscillatory for increasing K.



#### The Multitaper Method

The multitaper method controls the degrees of spectral *smoothing* and *averaging* through changing the properties of the tapers.

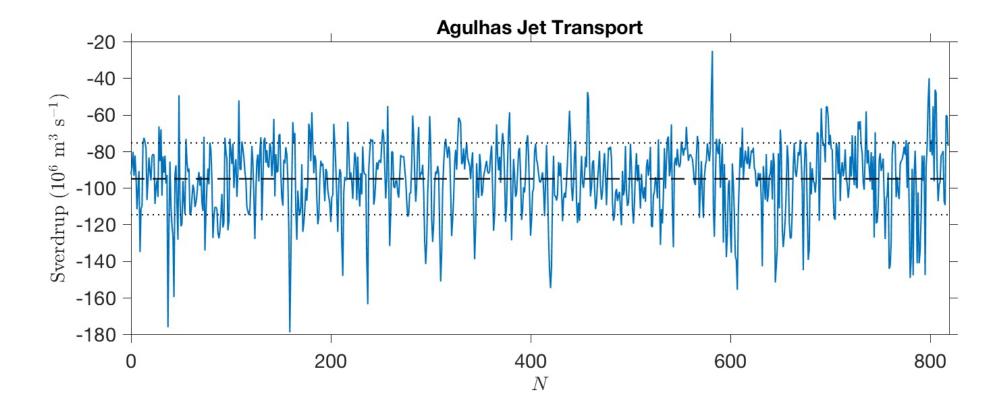
The multitaper method is generally the favorite spectral analysis method among those researchers who have thought the most about spectral analysis.

It is recommended because (i) it avoids the deficiencies of the periodogram, (ii) it has, in a certain sense, provable optimal properties, (iii) it is very easy to implement and adjust, (iv) it allows an estimate of the spectrum for the period equal to the length of your time series (no need to divide up your time series as for the Welch's method!).

See Thomson (1982), Park et al. (1987), and Percival and Walden, Spectral Analysis for Physical Applications.

## Example

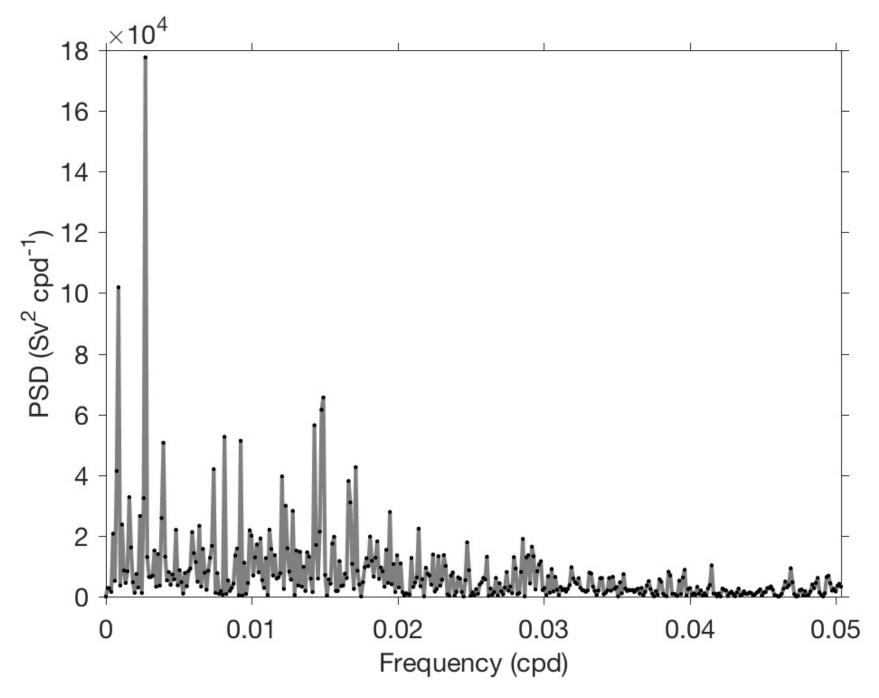
22-year time series of Agulhas current transport





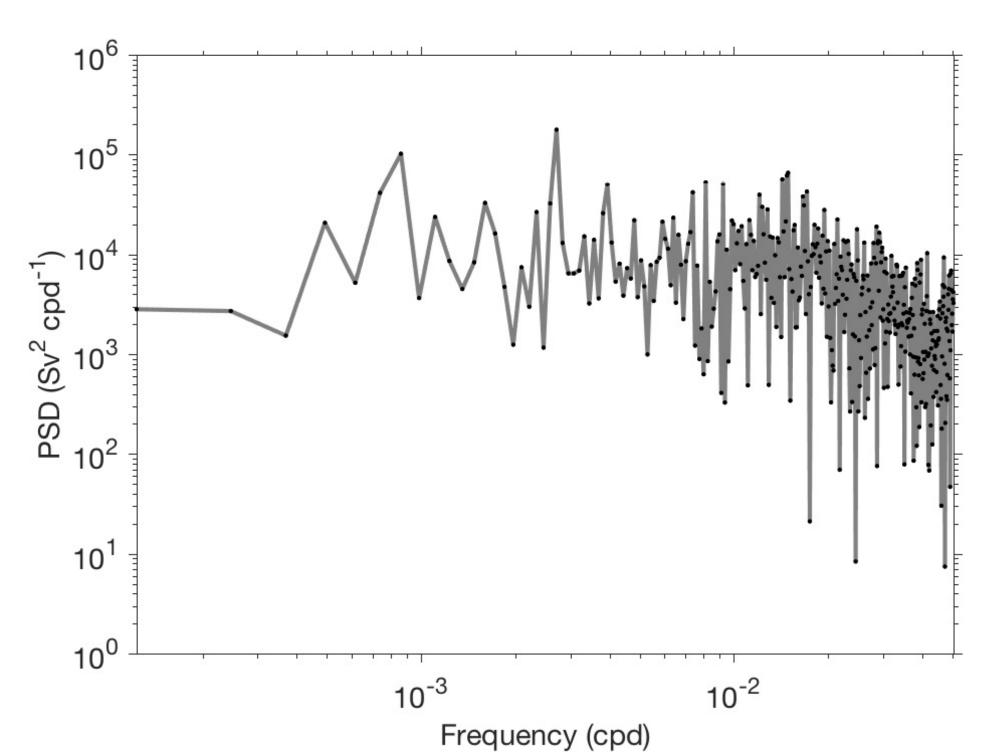
## Example: periodogram

#### Linear plot

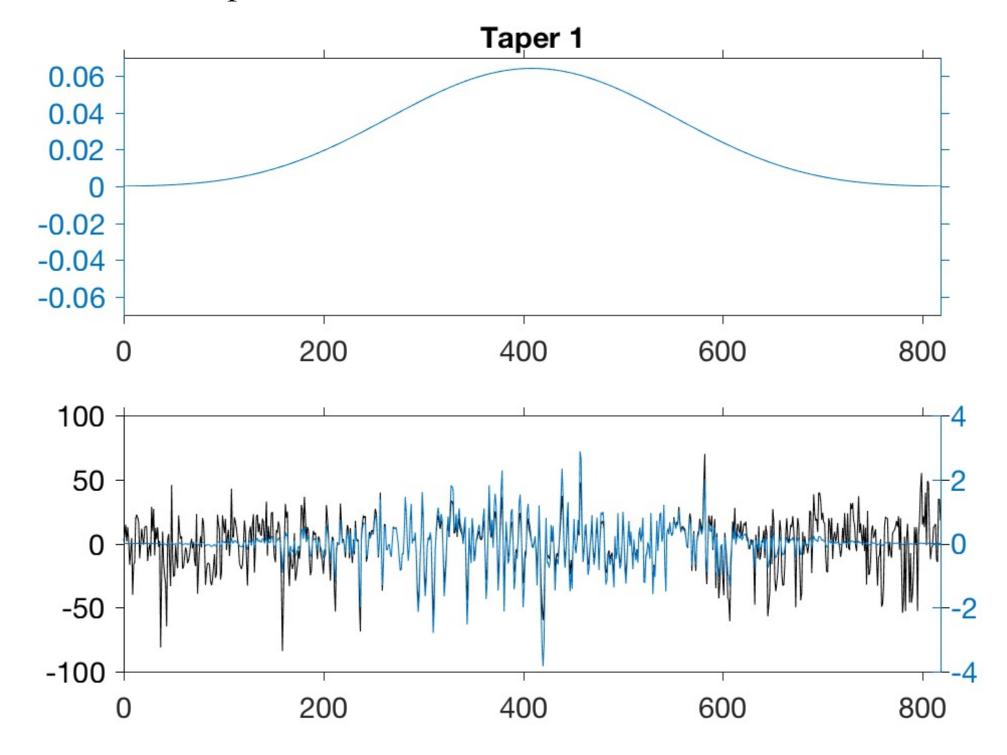


## Example: periodogram

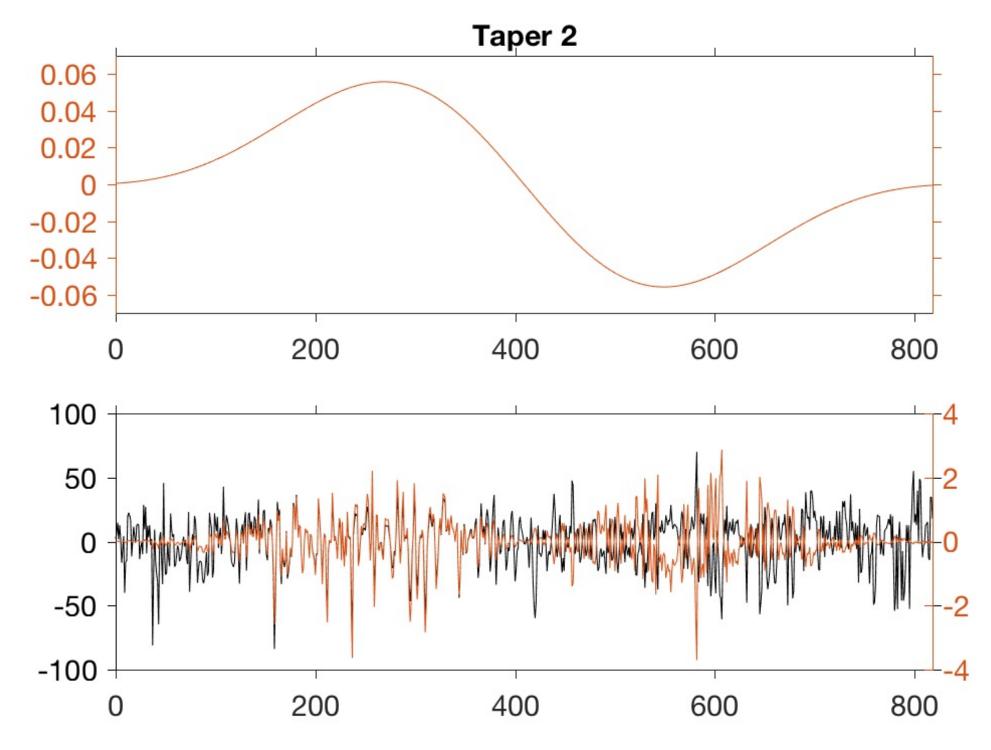
Log-log plot



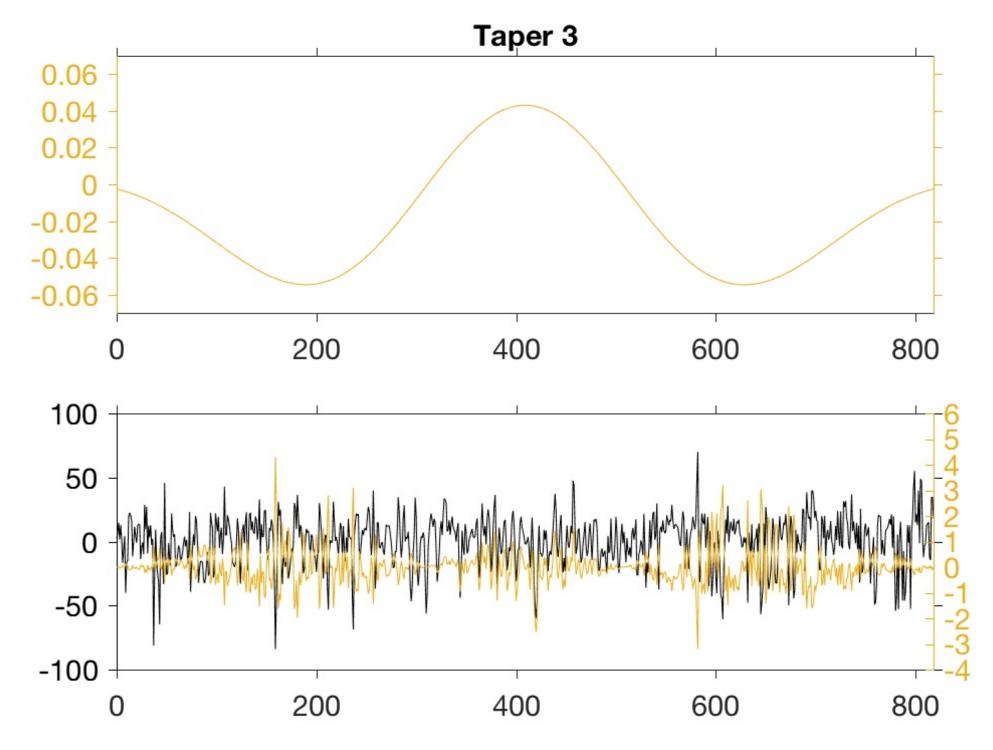
Effect of first taper



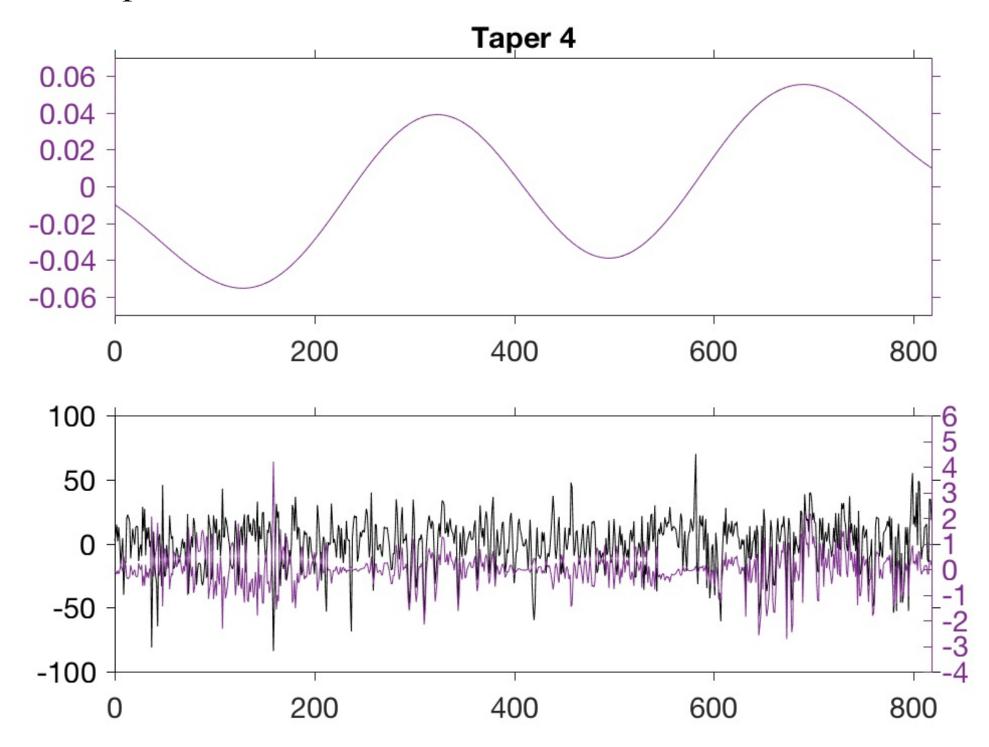
Second taper



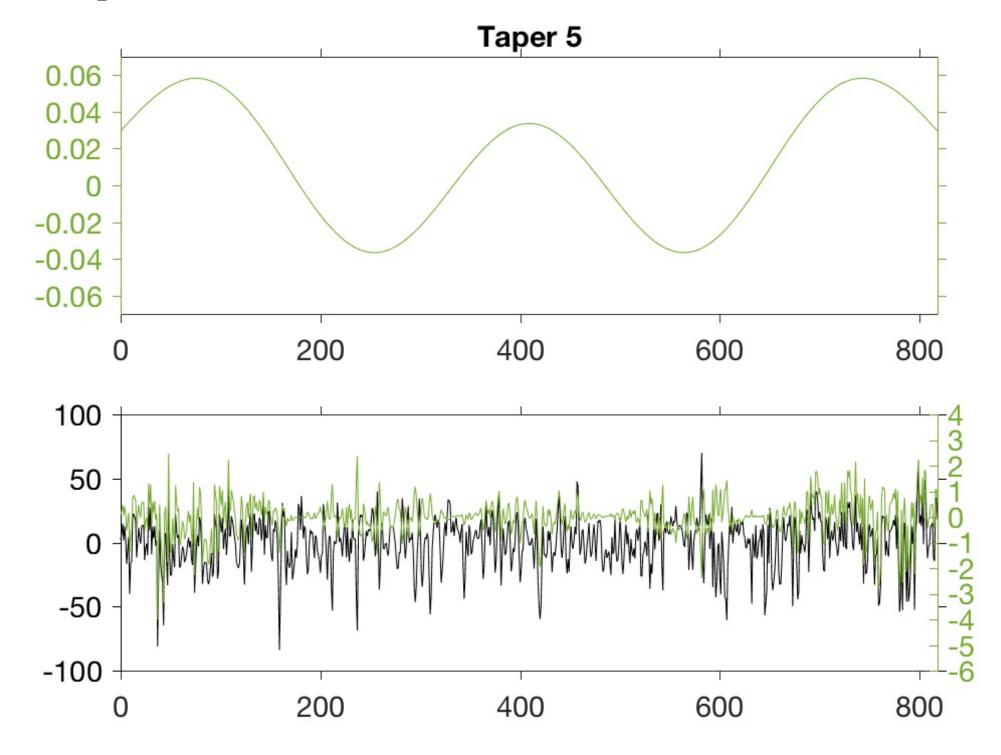
Third taper



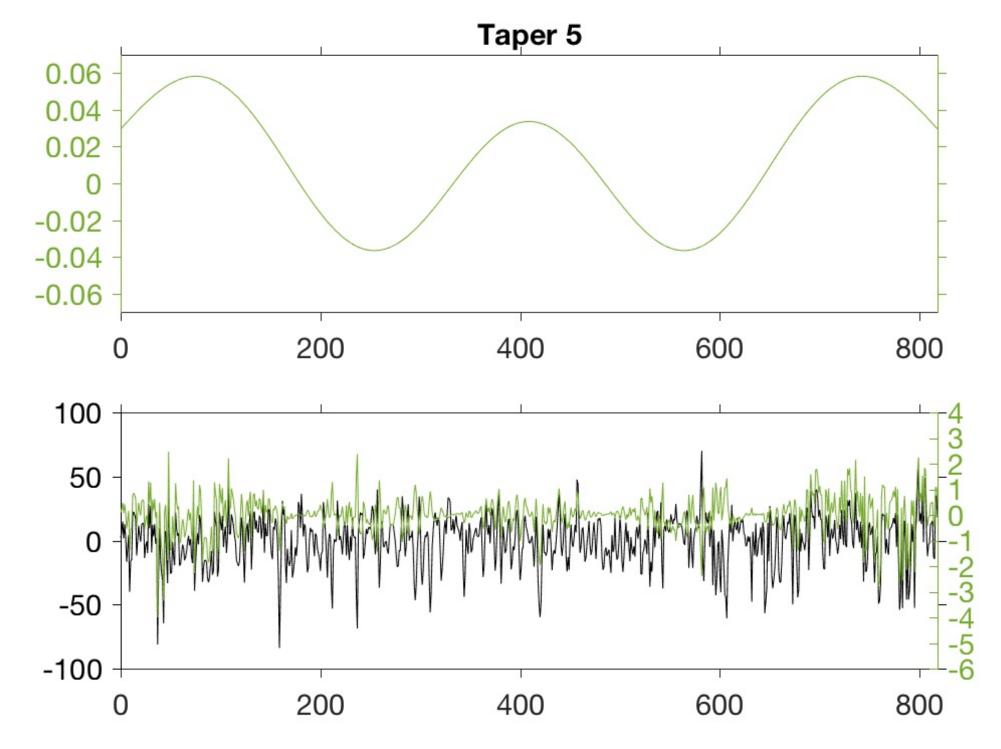
Fourth taper



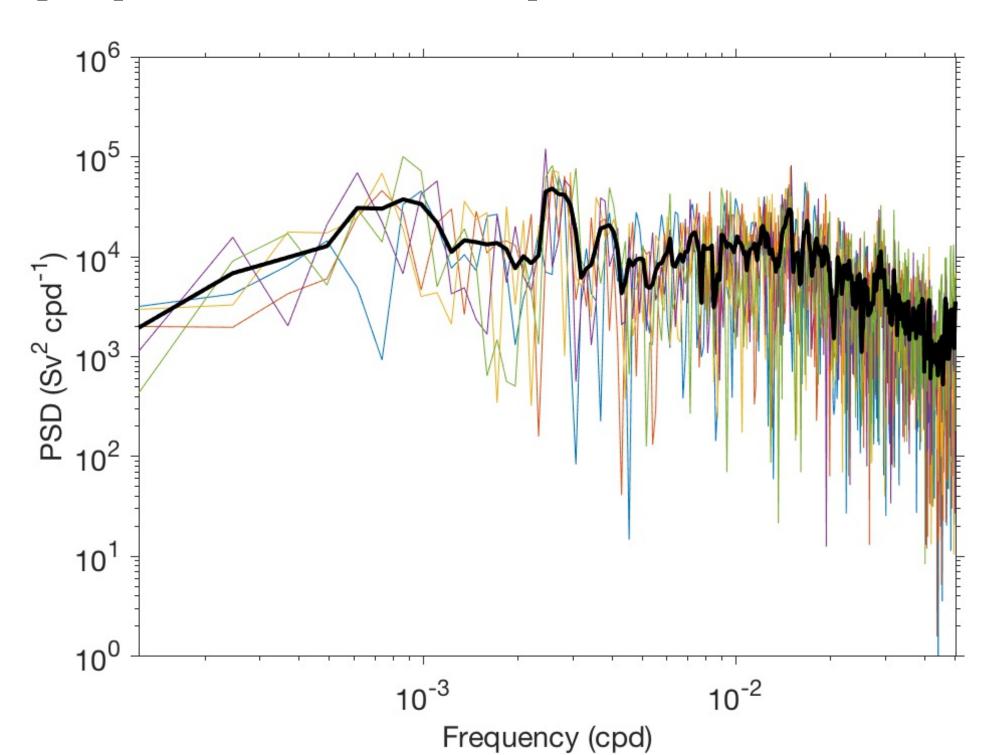
Fifth taper



Fifth taper

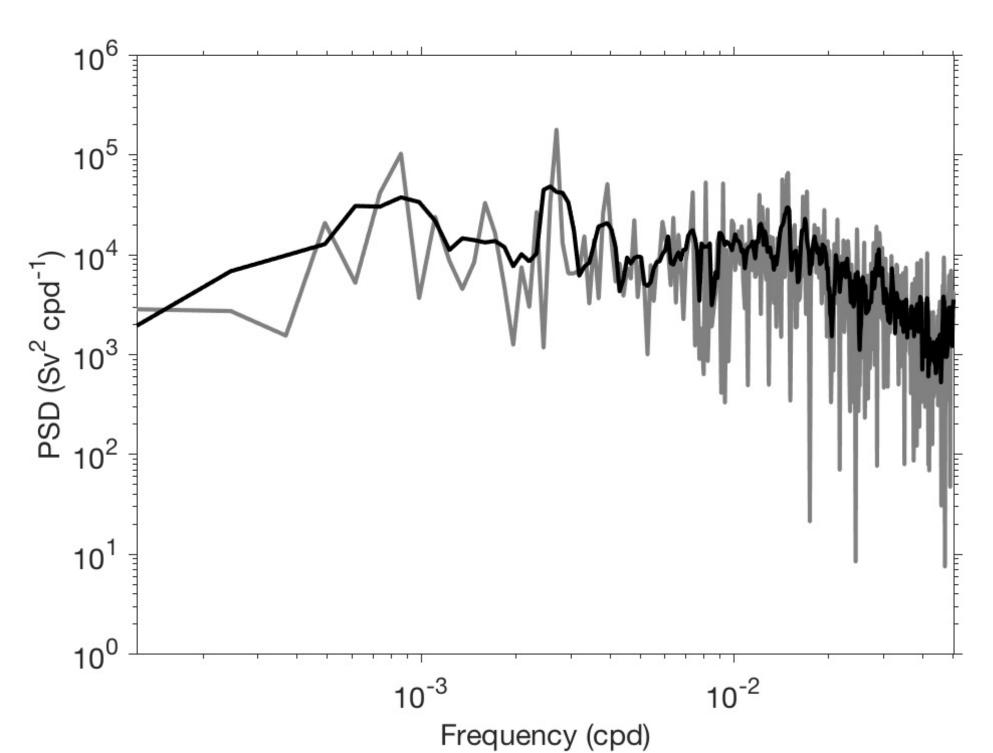


Eigen spectra (colors) and multitaper estimate (black)



## Example: period. vs mt

Periodogram (gray) vs multitaper (black)



## 4. Bivariate time series

# Vector and complex notations

What if your process of interest is composed of two time series, let's say x(t) and y(t)? As in the vector components of ocean currents or atmospheric winds:

$$\mathbf{z}(t) = egin{bmatrix} x(t) \ y(t) \end{bmatrix}$$

Often, a *bivariate* time series is conveniently written as a complex-valued time series:

$$z(t)=x(t)+iy(t)=|z(t)|e^{irg{(z)}},$$

where  $i = \sqrt{-1}$  and  $\arg(z)$  is the *complex argument* (or polar angle) of z in the interval  $[-\pi, +\pi]$ .

#### The Mean of Bivariate Data

The sample mean of the vector time series  $\mathbf{z}_n$  is also a vector,

$$\mu_{\mathbf{z}} \equiv rac{1}{N} \sum_{n=0}^{N-1} \mathbf{z}_n = \left[egin{array}{c} \hat{\mu}_x \ \hat{\mu}_y \end{array}
ight].$$

that consists of the *sample means* of the  $x_n$  and  $y_n$  components of  $\mathbf{z}_n$ .

#### Variance of Bivariate Data

The *variance* of the vector-valued times series  $\mathbf{z}_n$  is not a scalar or a vector, it is a  $2 \times 2$  *matrix* 

$$oldsymbol{\Sigma} \equiv rac{1}{N} \sum_{n=0}^{N-1} \left( \mathbf{z}_n - \mathbf{\mu}_{\mathbf{z}} 
ight) \left( \mathbf{z}_n - \mathbf{\mu}_{\mathbf{z}} 
ight)^T$$

where "T" is the  $matrix\ transpose,\ \mathbf{z}_n=\begin{bmatrix}x_n\\y_n\end{bmatrix},\ \mathbf{z}_n^T=[x_n\quad y_n]$  .

Carrying out the matrix multiplication leads to

$$oldsymbol{\Sigma} = rac{1}{N} \sum_{n=0}^{N-1} \left[ egin{array}{cc} (x_n - \hat{\mu}_x)^2 & (x_n - \hat{\mu}_x) \left(y_n - \hat{\mu}_y
ight) \ (x_n - \hat{\mu}_x) \left(y_n - \hat{\mu}_y
ight) & \left(y_n - \hat{\mu}_y
ight)^2 \end{array} 
ight]$$

The diagonal elements of  $\Sigma$  are the sample variances  $\sigma_x^2$  and  $\sigma_y^2$ , while the off-diagonal gives the *covariance* between  $x_n$  and  $y_n$ . Note that the two off-diagonal elements are identical.

#### Fourier transform

The Fourier theory presented earlier for scalar time series is completely applicable to complex-valued time series, in discrete form (N even):

$$egin{aligned} z_n &= rac{1}{N\Delta_t} \sum_{m=0}^{N/2} Z_m e^{i2\pi m n/N} \ &= rac{1}{N\Delta_t} \sum_{m=0}^{N-1} Z_m e^{i2\pi m n/N} \ &= rac{1}{N\Delta_t} \sum_{m=0}^{N/2} Z_m^+ e^{i2\pi m n/N} \ &+ rac{1}{N\Delta_t} \sum_{m=1}^{N/2-1} Z_m^- e^{-i2\pi m n/N}. \end{aligned}$$

The first sum corresponds to **positive** frequencies, and the second sum to the associated **negative** frequencies (except the zero and Nyquist frequencies for m = 0, N/2).

### Rotary Spectra

$$z_n = rac{1}{N\Delta_t} \sum_{m=0}^{N/2} Z_m^+ e^{i2\pi m n/N} + rac{1}{N\Delta_t} \sum_{m=1}^{N/2-1} Z_m^- e^{-i2\pi m n/N}.$$

This introduces the concept of *rotary spectrum*:

$$S(f_m>0)\equiv S^+(f_m)\equiv\lim_{N o\infty}E\left\{rac{|Z_m^+|^2}{N}
ight\} \quad ext{counterclockwise spectru} \ S(f_m<0)\equiv S^-(-f_m)\equiv\lim_{N o\infty}E\left\{rac{|Z_m^-|^2}{N}
ight\} \quad ext{clockwise spectrum.} \ f_m=rac{m}{N\Delta_t}, \quad m=0,\dots,N/2$$

This is very useful in geophysical fluid mechanics because counterclockwise motions are *cyclonic* in the northern hemisphere and clockwise motions are *anticyclonic*, and vice-versa in the southern hemisphere.

68 / 102

### Rotary variance

Imagine you have only two opposite components present in your time series at frequency  $f_k$ :

$$egin{align} z_n &= Z_k^+ e^{i2\pi k n/N} + Z_k^- e^{-i2\pi k n/N} \ &= \left\{A^+ e^{i\phi^+}
ight\} e^{i2\pi k n/N} + \left\{A^- e^{i\phi^-}
ight\} e^{-i2\pi k n/N} \ &= e^{i heta} \left\{A\cos(2\pi k n/N + \phi) + iB\sin(2\pi k n/N + \phi)
ight\} \end{aligned}$$

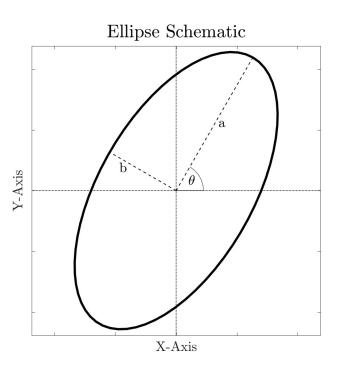
where

$$egin{aligned} heta &= (\phi^+ - \phi^-)/2 \ \phi &= (\phi^+ + \phi^-)/2 \ A &= A^+ + A^- \ B &= A^+ - A^-. \end{aligned}$$

### Elliptic variance

$$z_n = e^{i heta}\left\{A\cos(2\pi kn/N + \phi) + iB\sin(2\pi kn/N + \phi)
ight\}$$

This is the equation for an ellipse oriented at an angle  $\theta$  from the x axis, with semi-major and semi-minor axes A and B, respectively, rotating at frequency  $f_k = k/(N\Delta_t)$ , in the direction given by the sign of B.



See more details about elliptic variance in JML's Oslo lectures.

### Cartesian Spectra

Rotary and Cartesian spectra are two alternate representation of the variance of the complex time series:

$$z_n = rac{1}{N\Delta_t} \sum_{m=0}^{N/2} Z_m^+ e^{i2\pi m n/N} + rac{1}{N\Delta_t} \sum_{m=1}^{N/2-1} Z_m^- e^{-i2\pi m n/N}.$$

$${\widehat S}_m^+ \equiv rac{1}{N} {\left| Z_m^+ 
ight|^2}, \quad {\widehat S}_m^- \equiv rac{1}{N} {\left| Z_m^- 
ight|^2} \quad ext{Rotary spectra estimates}$$

$$z_n = x_n + i y_n = rac{1}{N\Delta_t} \sum_{m=0}^{N-1} X_m e^{i2\pi m n/N} + i \left\{ rac{1}{N\Delta_t} \sum_{m=0}^{N-1} Y_m e^{i2\pi m n/N} 
ight\}.$$

$$\widehat{S}_m^x \equiv rac{1}{N} |X_m|^2, \quad \widehat{S}_m^y \equiv rac{1}{N} |Y_m|^2 \quad ext{Cartesian spectra estimates}.$$

#### Parseval theorem

For bivariate data, the discrete form of the Parseval theorem takes the form:

$$egin{align} \Delta_t \sum_{n=0}^{N-1} |z_n|^2 &= rac{1}{N\Delta_t} \sum_{m=0}^{N-1} |Z_m|^2 &= rac{1}{N\Delta_t} \sum_{m=0}^{N-1} |X_m|^2 + rac{1}{N\Delta_t} \sum_{m=0}^{N-1} |Y_m|^2 \ &= rac{1}{N\Delta_t} \sum_{m=0}^{N/2} |Z_m^+|^2 + rac{1}{N\Delta_t} \sum_{m=1}^{N/2-1} |Z_m^-|^2 \ &= rac{1}{N\Delta_t} \sum_{m=0}^{N/2} |Z_m^+|^2 + rac{1}{N\Delta_t} \sum_{m=1}^{N/2-1} |Z_m^-|^2 \ &= rac{1}{N\Delta_t} \sum_{m=0}^{N/2-1} |Z_m^-|^2 \ &= \frac{1}{N\Delta_t} \sum_$$

#### Parseval theorem

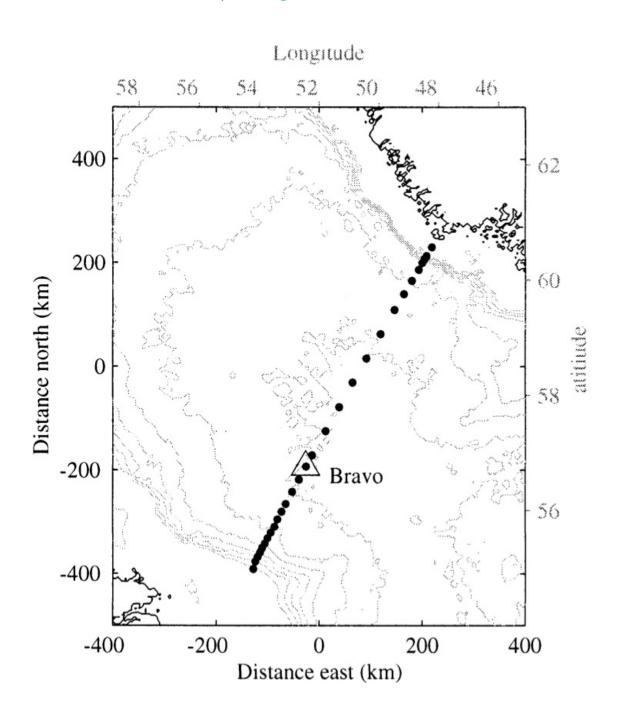
For bivariate data, the discrete form of the Parseval theorem takes the form:

$$egin{align} \Delta_t \sum_{n=0}^{N-1} |z_n|^2 &= rac{1}{N\Delta_t} \sum_{m=0}^{N-1} |Z_m|^2 &= rac{1}{N\Delta_t} \sum_{m=0}^{N-1} |X_m|^2 + rac{1}{N\Delta_t} \sum_{m=0}^{N-1} |Y_m|^2 \ &= rac{1}{N\Delta_t} \sum_{m=0}^{N/2} |Z_m^+|^2 + rac{1}{N\Delta_t} \sum_{m=1}^{N/2-1} |Z_m^-|^2 \ &= rac{1}{N\Delta_t} \sum_{m=0}^{N/2} |Z_m^+|^2 + rac{1}{N\Delta_t} \sum_{m=1}^{N/2-1} |Z_m^-|^2 \ &= rac{1}{N\Delta_t} \sum_{m=0}^{N/2-1} |Z_m^-|^2 \ &= \frac{1}{N\Delta_t} \sum_$$

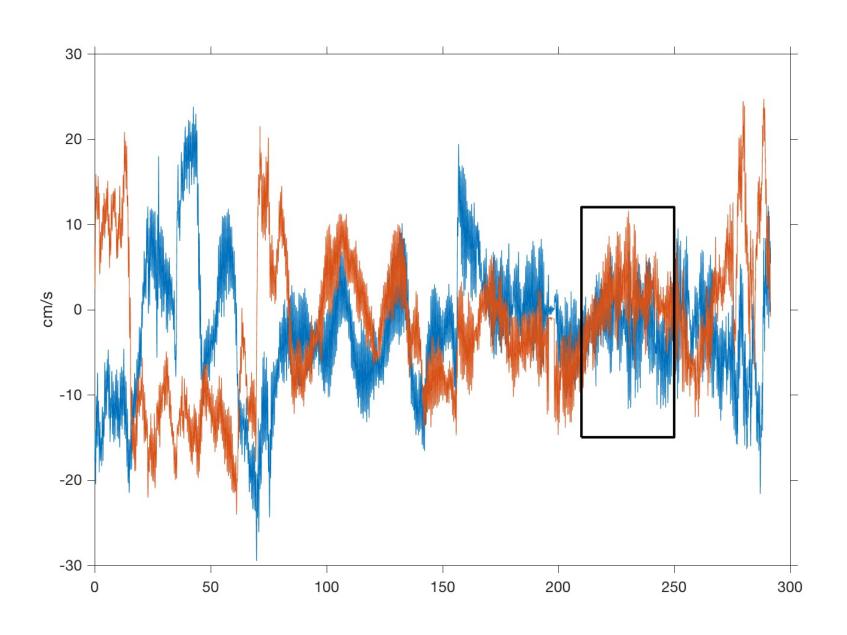
This shows that the total variance of the bivariate process is recovered completely by the Cartesian, or rotary Fourier representation.



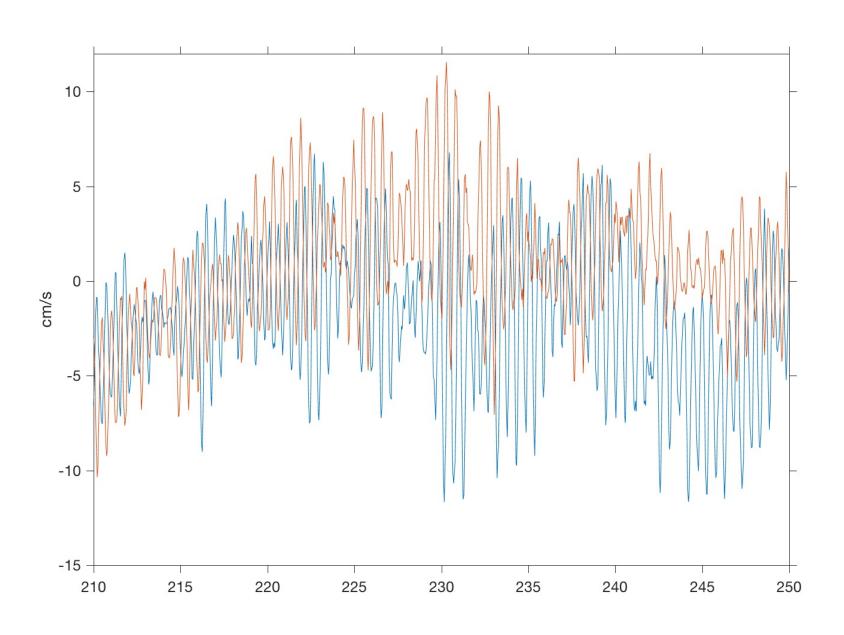
Hourly current meter record at 110 m depth from the Bravo mooring in the Labrador Sea, *Lilly and Rhines* (2002)



Hourly current meter record at 110 m depth from the Bravo mooring in the Labrador Sea, *Lilly and Rhines* (2002)



Hourly current meter record at 110 m depth from the Bravo mooring in the Labrador Sea, *Lilly and Rhines* (2002)



#### Observable Features

- 1. The data consists of two time series that are similar in character.
- 2. Both time series present a superposition of scales.
- 3. At the smallest scale, there is an apparently oscillatory roughness which changes its amplitude in time.
- 4. A larger scale presents itself either as localized features, or as wavelike in nature.
- 5. Several sudden transitions are associated with isolated events.
- 6. Zooming in, we see the small-scale oscillatory behavior is sometimes  $90^{\circ}$  degrees out of phase, and sometimes  $180^{\circ}$ .
- 7. The amplitude of this oscillatory variability changes with time.

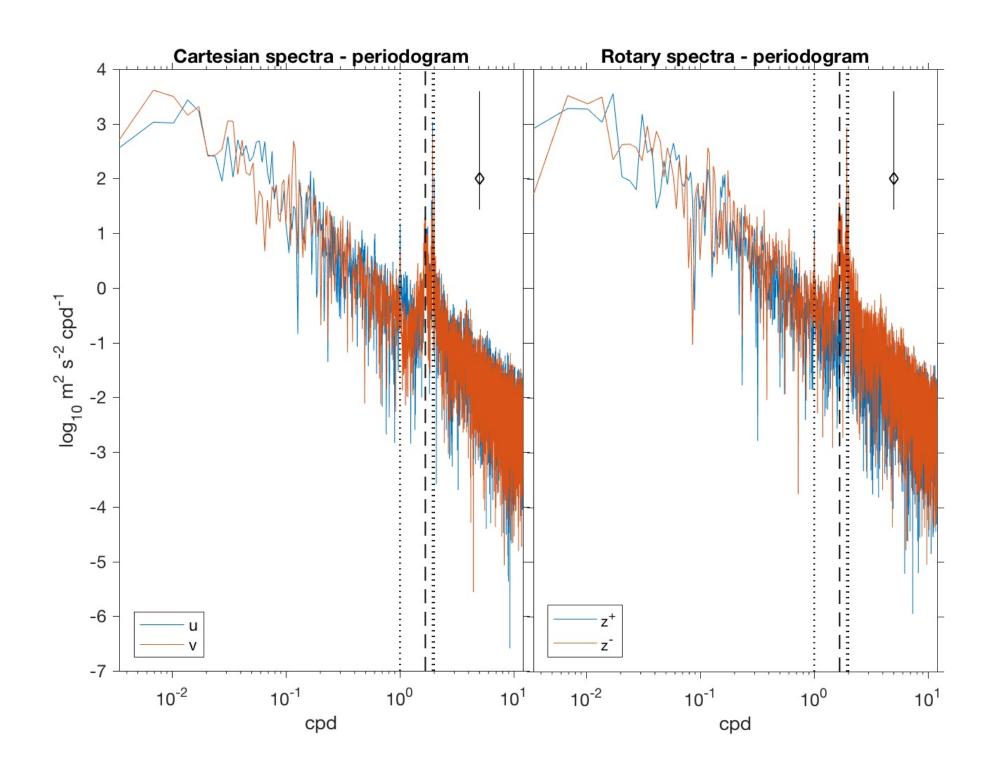
The fact that the oscillatory behavior is not consistently  $90^{\circ}$  out of phase removes the possibility of these features being purely inertial oscillations. The amplitude modulation suggests tidal beating.

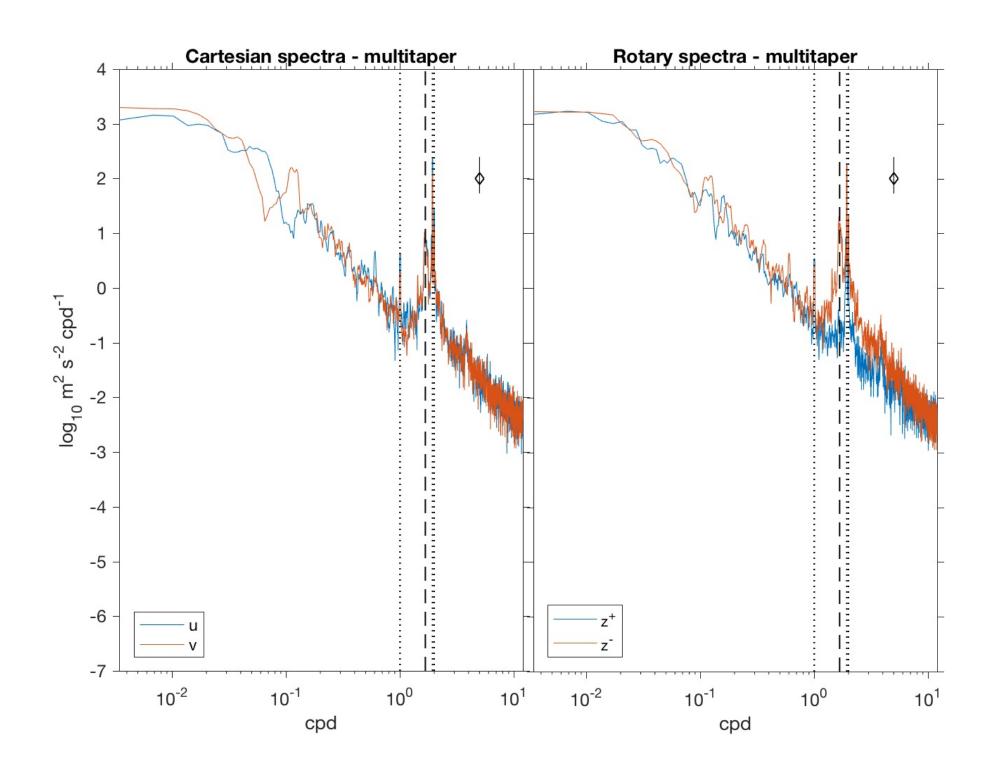
#### Observable Features

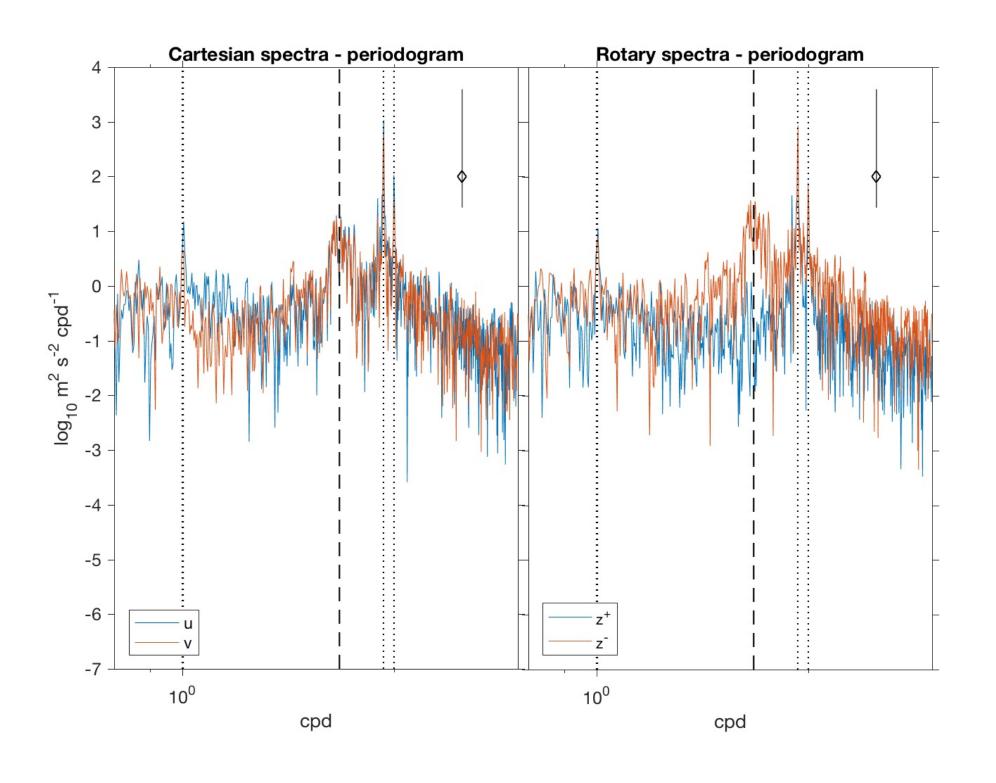
- 1. The data consists of two time series that are similar in character.
- 2. Both time series present a superposition of scales.
- 3. At the smallest scale, there is an apparently oscillatory roughness which changes its amplitude in time.
- 4. A larger scale presents itself either as localized features, or as wavelike in nature.
- 5. Several sudden transitions are associated with isolated events.
- 6. Zooming in, we see the small-scale oscillatory behavior is sometimes  $90^{\circ}$  degrees out of phase, and sometimes  $180^{\circ}$ .
- 7. The amplitude of this oscillatory variability changes with time.

The fact that the oscillatory behavior is not consistently  $90^{\circ}$  out of phase removes the possibility of these features being purely inertial oscillations. The amplitude modulation suggests tidal beating.

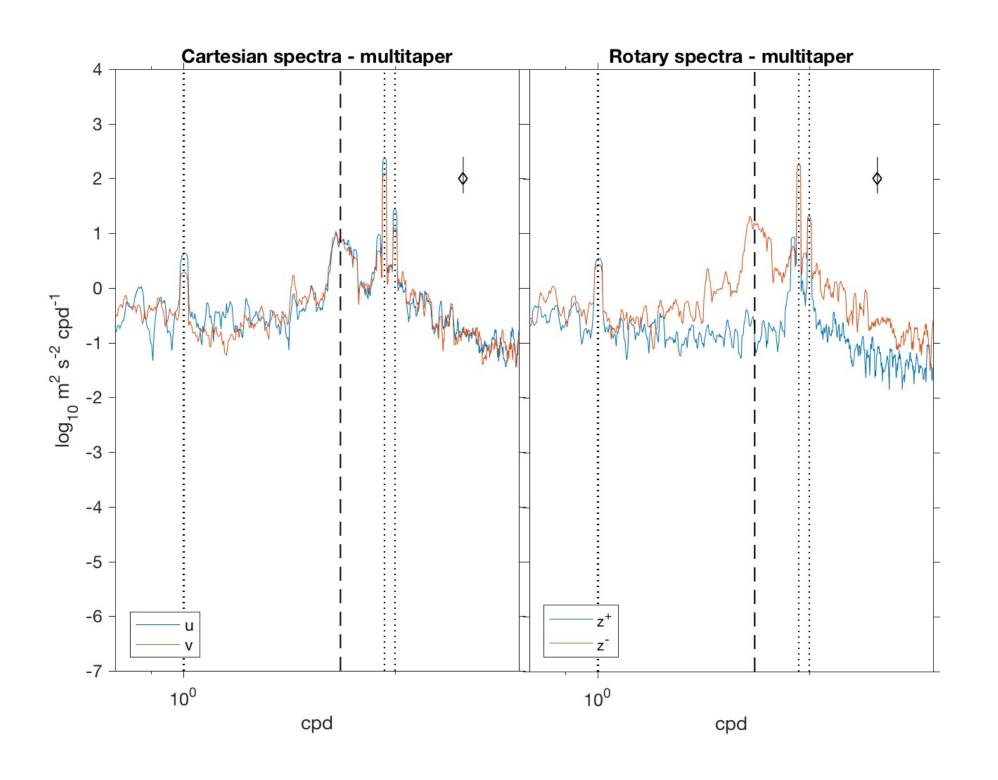
The isolated events are eddies, which cause the currents to suddenly rotate as they pass by. The oscillations are due to tides and internal waves.





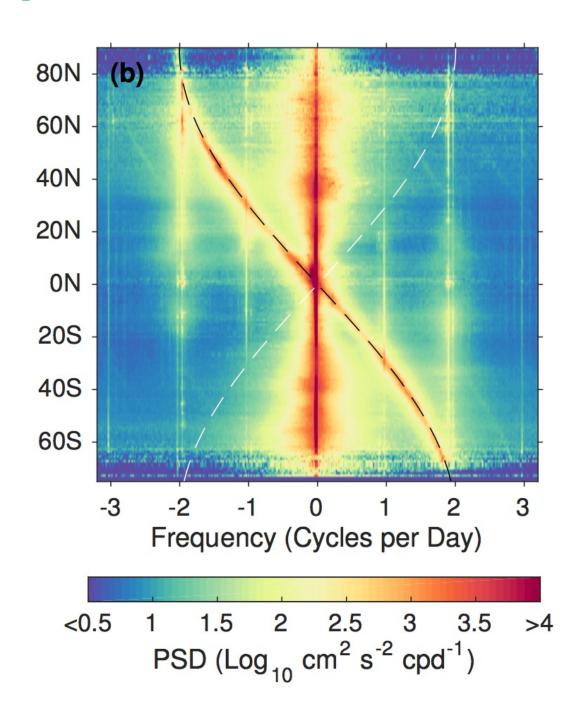








Global zonally-averaged rotary spectra from hourly drifter velocities, see *Elipot et al. 2016*.



### **Epilogue**

Many important topic of time series analyses were not discussed, namely the topics of *aliasing*, spectral blurring, filtering, and uncertainties in spectral estimates. Some of these are covered in JML's Oslo lecture which can be viewed online at {www.jmlilly.net/talks/oslo/index.html}.

During the practical session this afternoon we will cover the material presented this morning, as well cover some of the topic of filtering.

Thank you!

Shane Elipot

email: selipot@rsmas.miami.edu

The slides of this presentation are available at {rsmas.miami.edu/users/selipot/} or {selipot.github.io}



### Slides for Pratical session



#### Code and Data

For the practical session this afternoon you will need some Matlab code and data to be downloaded from my GitHUb depository called *Tromso*.

https://github.com/selipot/tromso.

Click on the green button "Clone or download" to download a zipped archive, expand it, and copy the directory where you want.

Please do this before coming to the session this afternoon.



#### Continuous Fourier

We have considered the FT for a discrete time series  $x_n$  but it extends to continuous time series x(t):

$$x(t) = rac{1}{2\pi} \int_{-\infty}^{\infty} X(\omega) \, e^{i\omega t} d\omega, \qquad X(\omega) \equiv \int_{-\infty}^{\infty} x(t) \, e^{-i\omega t} dt.$$

Note that here we are using radian frequency  $\omega$ .

#### Continuous Fourier

We have considered the FT for a discrete time series  $x_n$  but it extends to continuous time series x(t):

$$x(t)=rac{1}{2\pi}\int_{-\infty}^{\infty}X(\omega)\,e^{i\omega t}d\omega, \qquad X(\omega)\equiv\int_{-\infty}^{\infty}x(t)\,e^{-i\omega t}dt.$$

Note that here we are using radian frequency  $\omega$ .

The continuous notation is easier to understand the mechanics of *filtering* a time series, as well as spectral *blurring*.



#### Continuous Fourier

We have considered the FT for a discrete time series  $x_n$  but it extends to continuous time series x(t):

$$x(t)=rac{1}{2\pi}\int_{-\infty}^{\infty}X(\omega)\,e^{i\omega t}d\omega, \qquad X(\omega)\equiv\int_{-\infty}^{\infty}x(t)\,e^{-i\omega t}dt.$$

Note that here we are using radian frequency  $\omega$ .

The continuous notation is easier to understand the mechanics of *filtering* a time series, as well as spectral *blurring*.

But first, we need to understand the concept of convolution.

### The Convolution Integral

The convolution h(t) of a function f(t) and g(t) is defined as:

$$h(t) \equiv \int_{-\infty}^{\infty} f( au) g(t- au) d au.$$

Note that in convolution, the order does not matter and we can show that

$$h(t) \equiv \int_{-\infty}^{\infty} g( au) f(t- au) d au$$

This mathematical operation is actually what is being done when "smoothing" data. (It is like sliding the iron on the tablecloth, or pulling the tablecloth under a static iron).

#### **Convolution Theorem**

The convolution theorem states convolving f(t) and g(t) in the time domain is the same as a multiplication in the frequency domain.

Let  $F(\omega)$  and  $G(\omega)$  be the Fourier transforms of f(t) and g(t), respectively. It can be shown that if

$$h(t) = \int_{-\infty}^{\infty} f( au) g(t- au) d au$$

then the fourier transform of h(t) is

$$H(\omega) = F(\omega)G(\omega).$$

#### **Convolution Theorem**

The convolution theorem states convolving f(t) and g(t) in the time domain is the same as a multiplication in the frequency domain.

Let  $F(\omega)$  and  $G(\omega)$  be the Fourier transforms of f(t) and g(t), respectively. It can be shown that if

$$h(t) = \int_{-\infty}^{\infty} f( au) g(t- au) d au$$

then the fourier transform of h(t) is

$$H(\omega)=F(\omega)G(\omega).$$

This result is key to understand what happens in the Fourier domain when you perform a time-domain smoothing.

### Smoothing

Now we consider what happens when we *smooth* the time series x(t) by the filter g(t) to obtain a smoothed version  $\tilde{x}(t)$  of your time series:

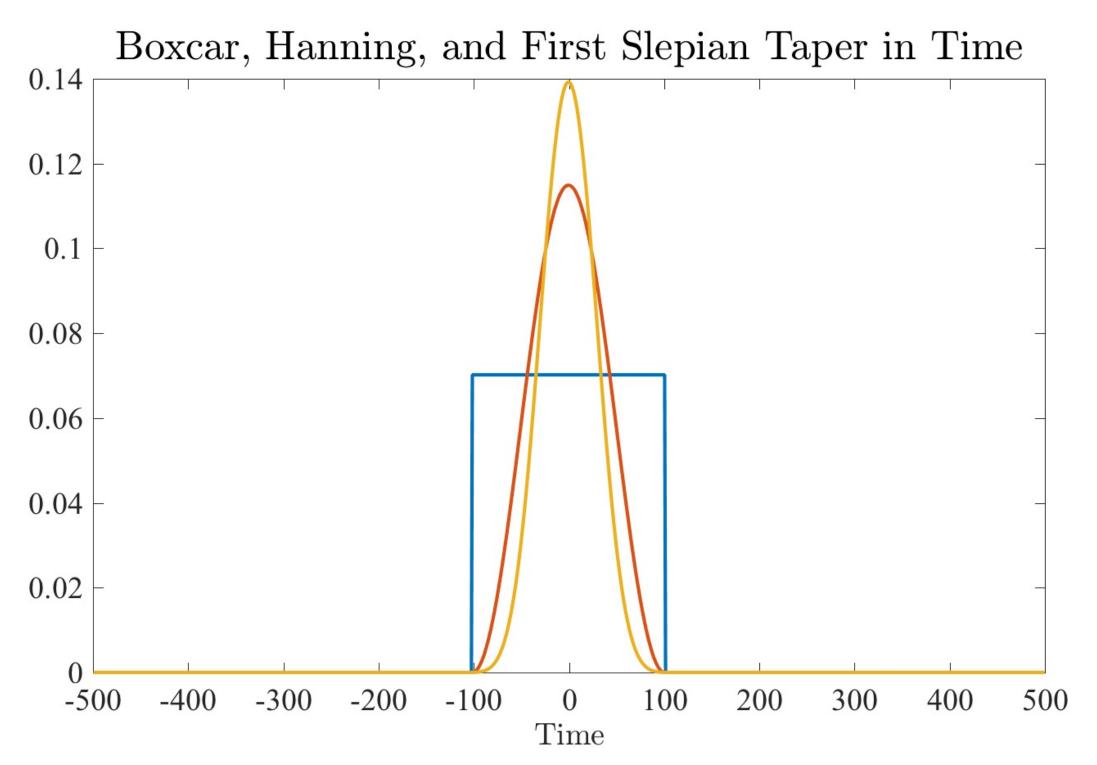
$$egin{aligned} ilde{x}(t) &= \int_{-\infty}^{\infty} x(t- au) g( au) d au \equiv rac{1}{2\pi} \int_{-\infty}^{\infty} \widetilde{X}(\omega) \, e^{i\omega t} d\omega. \ &= rac{1}{2\pi} \int_{-\infty}^{\infty} X(\omega) G(\omega) \, e^{i\omega t} d\omega, \end{aligned}$$

by the convolution theorem.

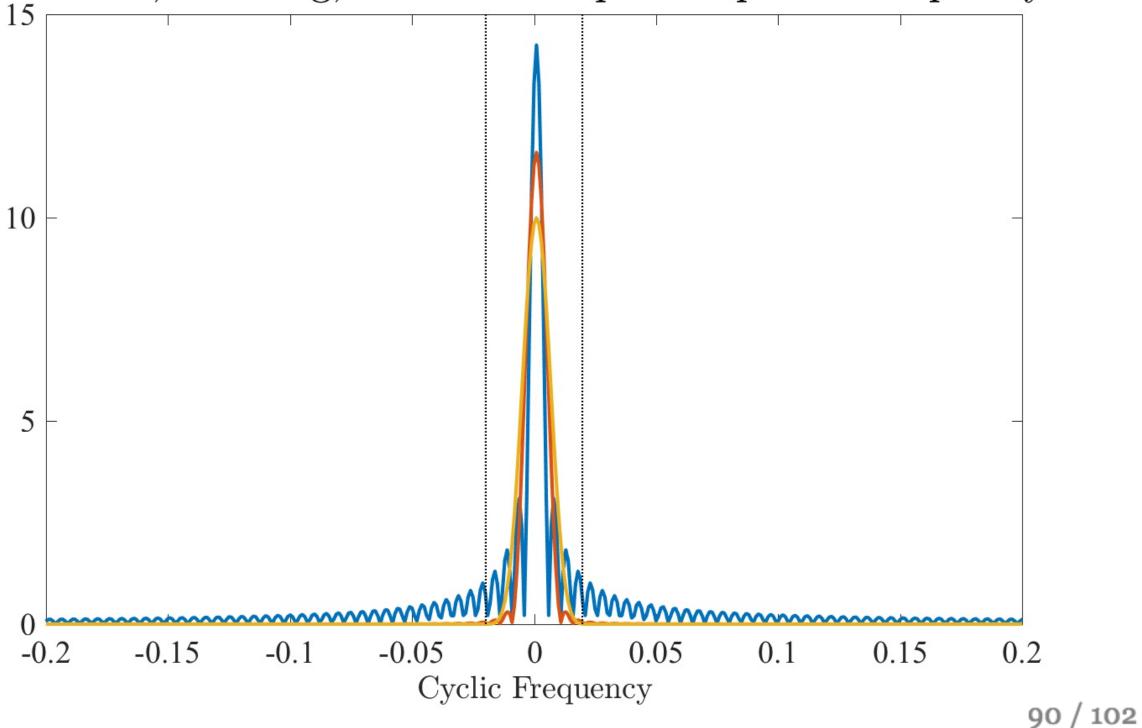
When we perform simple smoothing, we are also reshaping the Fourier transform of the signal by *multiplying* its Fourier transform by that of the smoothing window.



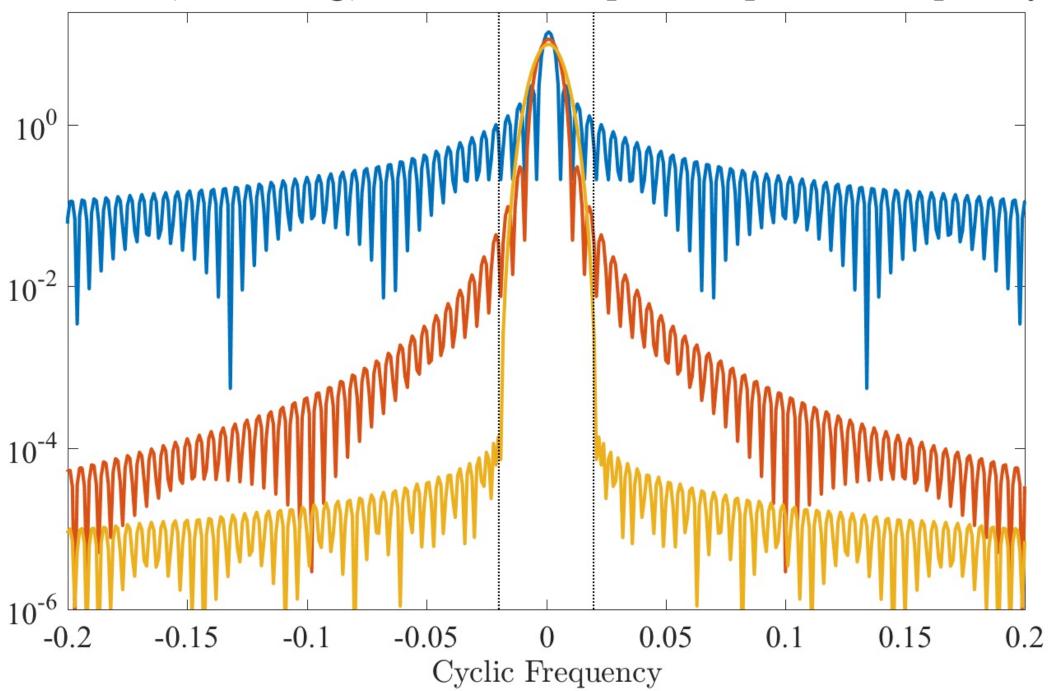
### Three Window examples



Boxcar, Hanning, and First Slepian Taper in Frequency



Boxcar, Hanning, and First Slepian Taper in Frequency



### Lowpass & Highpass Filters

From the convolution theorem, we understand that filtering will keep the frequencies near zero but reject higher frequencies. For this reason they are called *low-pass filters*.

The reverse type of filtration, rejecting the low frequencies but keeping the high frequencies, is called *high-pass filtering*.

The residual  $\check{x}(t) \equiv x(t) - \tilde{x}(t)$  is an example of a high-pass filtered time series.

In practice, to find the frequency form of your filter, you *pad it with zeros* so that it becomes the same length as your time series, and then you take its discrete Fourier transform.



# Back to practical!

#### **Convolution Theorem II**

This theorem is reciprocal: is your *multiply* in the time domain, you *convolve* in the Fourier domain.

It can be shown that if

$$h(t) = f(t)g(t)$$

then the fourier transform of h(t) is

$$H(\omega) = \int_{-\infty}^{\infty} F(
u) G(\omega - 
u) d\omega.$$

This result is key to understand what happens in the Fourier domain when you try to estimate spectra, i.e. *spectral bluring*, or to design band-pass filters.

### Bandpass filtering

We can use the convolution theorem to build a band-pass filter.

We want to modify the lowpass filter g(t) so that its Fourier transform is localized not about zero, but about some non-zero frequency  $\omega_o$ . To do this, we multiply g(t) by a complex exponential  $g(t)e^{i\omega_o t}$ .

It can be shown (see Oslo lectures) that the Fourier transform of  $g(t)e^{i\omega_o t}$  is  $G(\omega-\omega_o)$ , which is localized around  $\omega_o$ .

Thus, a convolution with  $g(t)e^{i\omega_o t}$  will bandpass the data in the vicinity of  $\omega_o$ .

In fact, a lowpass filter is a particular type of bandpass in which the center of the *pass band* has been chosen as zero frequency.

# Back to practical!



#### **Effect of Truncation**

Now imagine instead that we have a *continuously* sampled time series of length T, that is, we have z(t) but only between times -T/2 and T/2. This is like multiplying z(t) by a function g(t) which is equal to one between -T/2 and T/2 and o otherwise:

$$z_T(t) = g(t) imes z(t)$$

We will denote this truncted version of z(t) by  $z_T(t)$ .

How does the spectrum compare of  $z_T(t)$  compare with that of z(t)?

#### Effect of Truncation

Now imagine instead that we have a *continuously* sampled time series of length T, that is, we have z(t) but only between times -T/2 and T/2. This is like multiplying z(t) by a function g(t) which is equal to one between -T/2 and T/2 and o otherwise:

$$z_T(t) = g(t) imes z(t)$$

We will denote this truncted version of z(t) by  $z_T(t)$ .

How does the spectrum compare of  $z_T(t)$  compare with that of z(t)?

Q. Using one of the windows encountered today, how can we express the relationship between z(t) and  $z_T(t)$ ?

#### Effect of Truncation

Now imagine instead that we have a *continuously* sampled time series of length T, that is, we have z(t) but only between times -T/2 and T/2. This is like multiplying z(t) by a function g(t) which is equal to one between -T/2 and T/2 and o otherwise:

$$z_T(t) = g(t) imes z(t)$$

We will denote this truncted version of z(t) by  $z_T(t)$ .

How does the spectrum compare of  $z_T(t)$  compare with that of z(t)?

Q. Using one of the windows encountered today, how can we express the relationship between z(t) and  $z_T(t)$ ?

Q. Therefore, using another theorem learned today, what is the difference between their spectra?

### Spectral Blurring

The spectrum of the truncated time series is *blurred* through smoothing with a function  $F_T(\omega)$  that is the *square* of the Fourier transform of a boxcar:

$$\widetilde{S}(\omega) \equiv rac{1}{2\pi} \int_{-\infty}^{\infty} S(
u) \, F_T(
u - \omega) \, d
u.$$

The smoothing function, which is known as the Fejér kernel

$$F_T(\omega) \equiv \int_{-T}^T \left(1-rac{| au|}{T}
ight)\,e^{i\omega au}\,d au = rac{1}{T}rac{\sin^2(\omega T/2)}{(\omega/2)^2}\,.$$

is essentially a squared version of the "sinc" or  $\sin(x)/x$  function.

However,  $\sin(x)/x$  is not a very smooth function at all!

### Multitapering Revisited

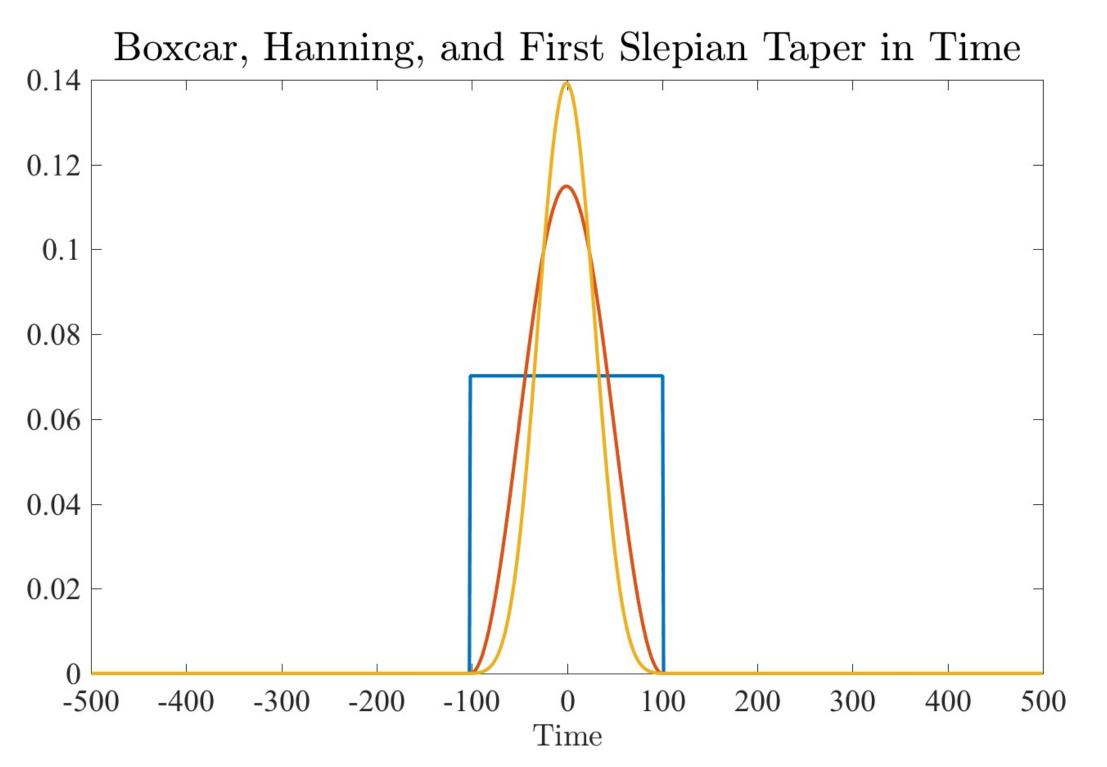
We can now understand the purpose of multitapering.

Doing nothing in your spectral estimate is equivalent to truncating your data, thus implicitly smoothing the true spectrum by an extremely undesirable function!

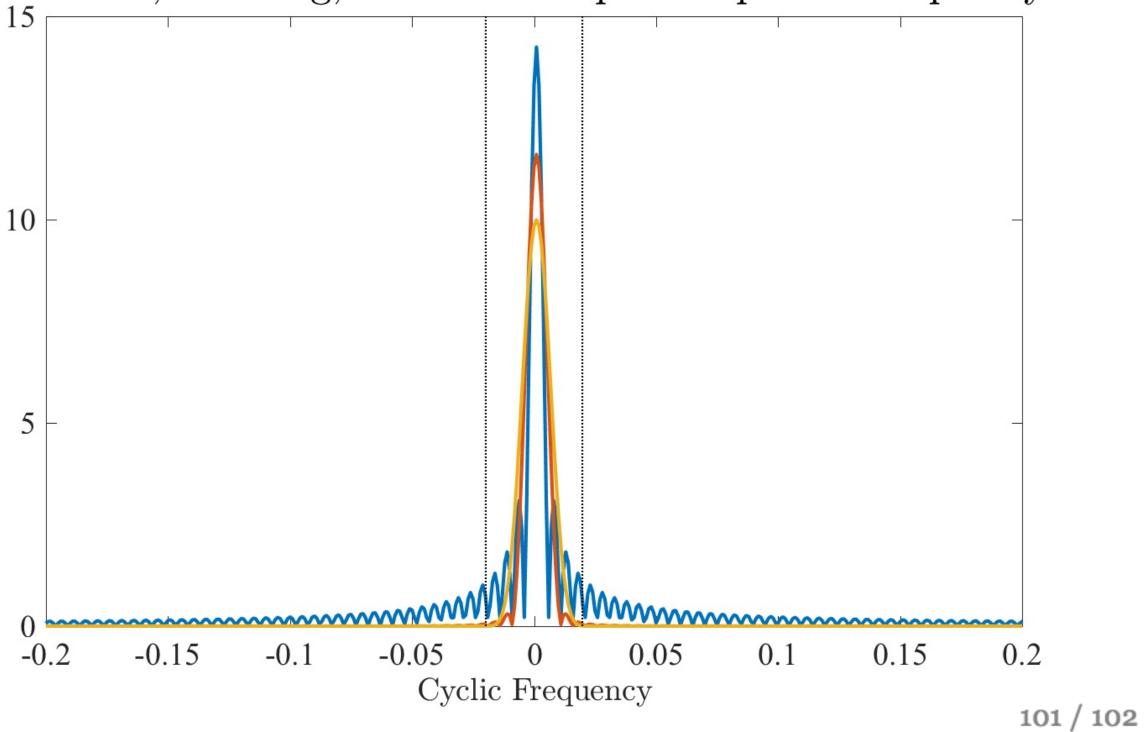
The Fejér kernel has a major problem in that it is not *well* concentrated. Its "side lobes" are large, leading to a kind of error called *broadband bias*.

This is the source of the error shown in the motivating example.

Next we take a look at three different tapering functions and their squared Fourier transforms. The broadband bias is most clear if we use logarithmic scaling for the y-axis.



Boxcar, Hanning, and First Slepian Taper in Frequency



Boxcar, Hanning, and First Slepian Taper in Frequency

